

Flux vacua of the mirror octic

Erik Plauschinn

Utrecht University

Geometry, Strings, and the Swampland

Ringberg Castle — 19.03.2024

based on

This talk is based on ::

Flux vacua of the mirror octic

E. Plauschinn, L. Schlechter

arXiv:2310.06040

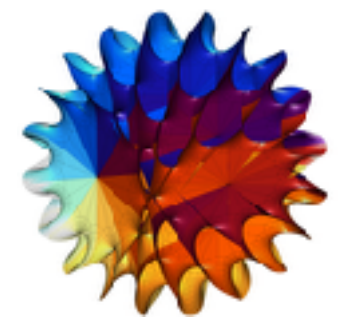
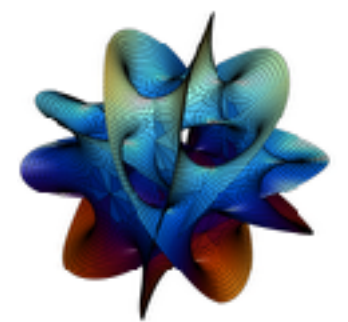
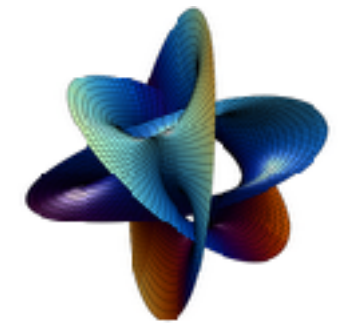
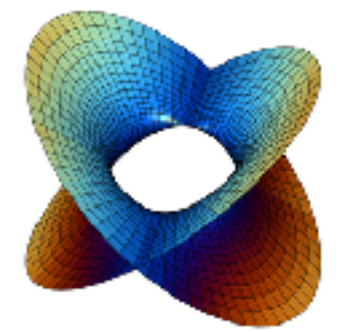
JHEP 01 (2024) 157

Some motivation.

Compactifications of string theory on Calabi-Yau three-folds give rise to four-dimensional effective theories.

These theories often contain massless scalar fields (moduli) that parametrize the compact space.

Deforming the compact space by fluxes can make moduli massive.

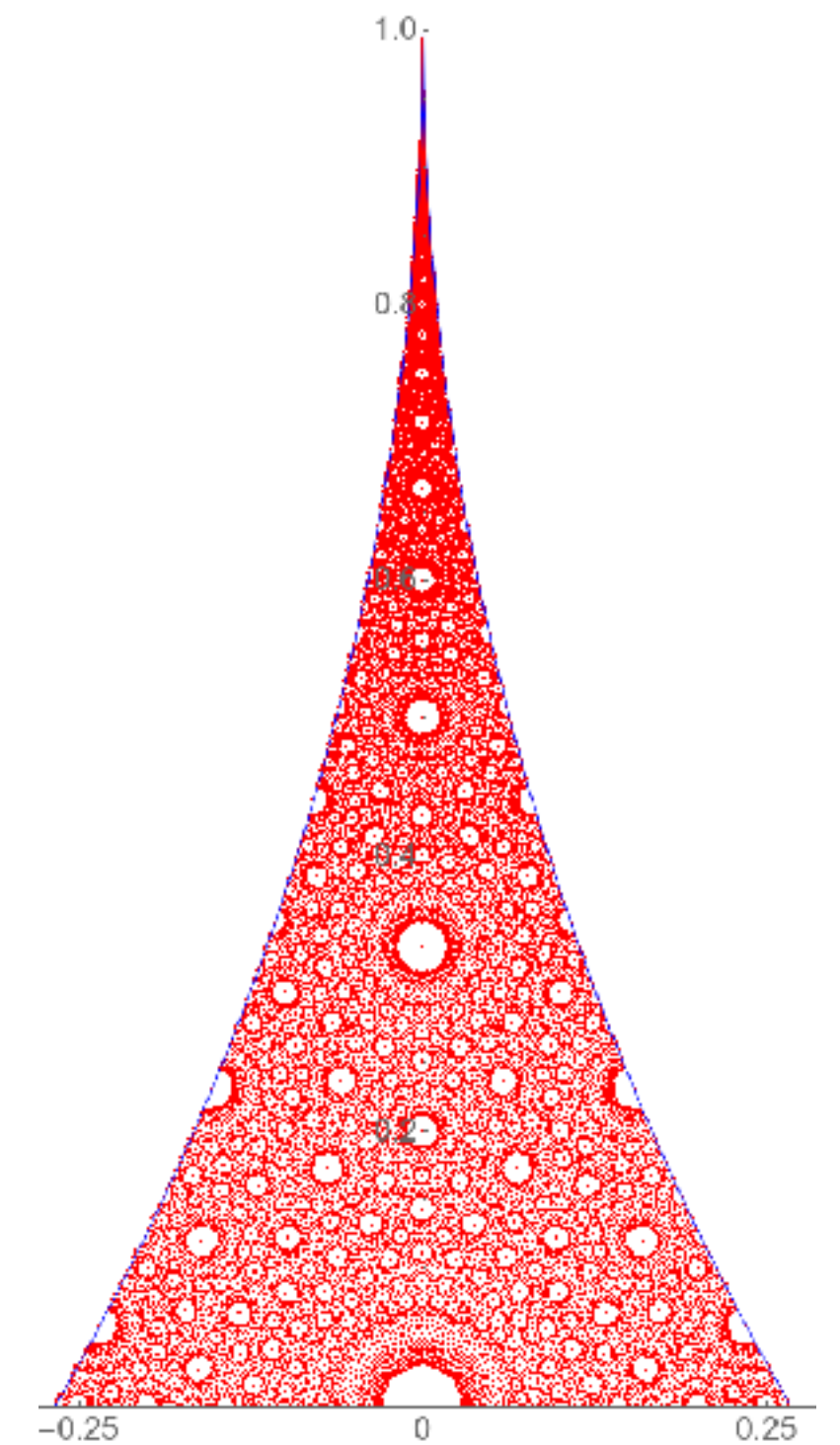


In type IIB orientifold compactifications on Calabi-Yau three-folds $\mathcal{X}$, **fluxes** generate mass-terms for the axio-dilaton and complex-structure moduli.

Dasgupta, Rajesh, Sethi — 1999
Giddings, Kachru, Polchinski — 2001

The fluxes are constrained by the **tadpole cancellation condition**

$$0 < N_{\text{flux}} \leq N_{\text{max}} , \quad N_{\text{flux}} = \int_{\mathcal{X}} F \wedge H .$$



Questions ::

- Is the **number of flux vacua** for a given $N_{\max}$ **finite**?

Grimm — 2020

Bakker, Grimm, Schnell, Tsimmerman — 2021

- **How many** flux vacua exist for a given $N_{\max}$?

Ashok, Douglas — 2003

Denef, Douglas — 2004

- What are general **properties** of such flux vacua?

DeWolfe, Giryavets, Kachru, Taylor — 2004

Conlon, Quevedo — 2004

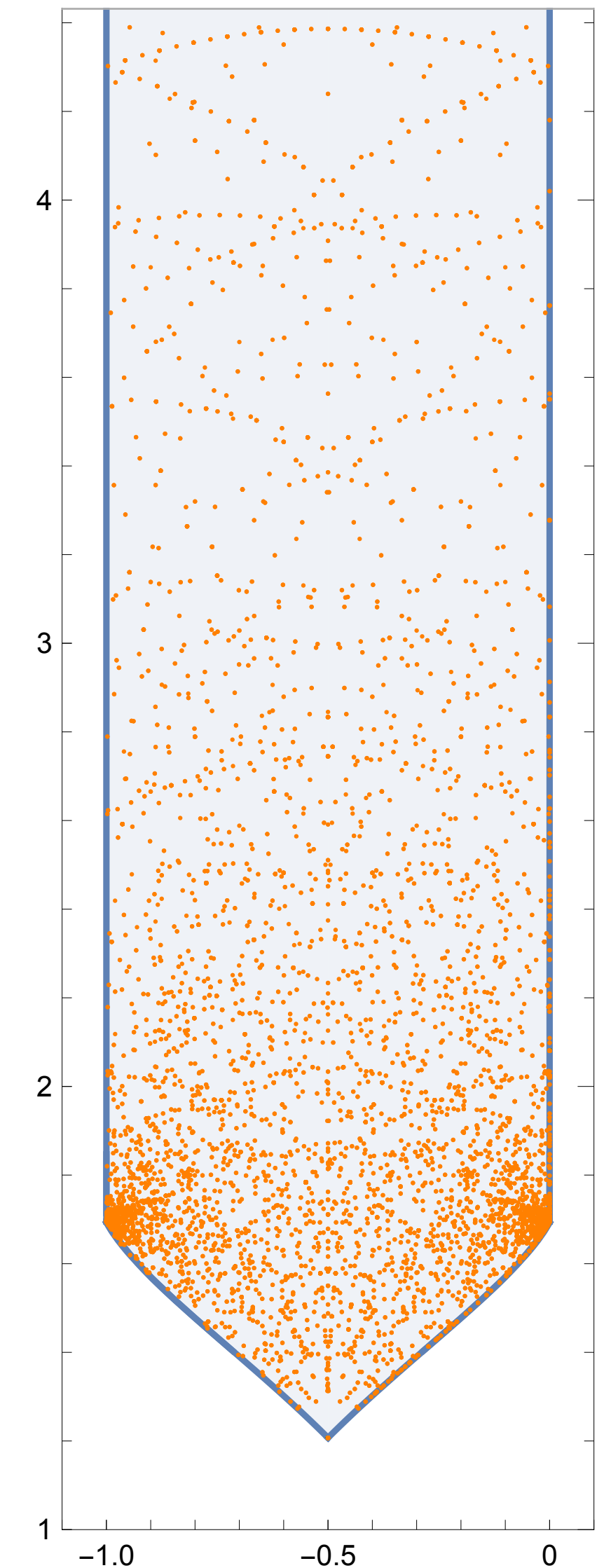
Cole, Shiu — 2018

Dubey, Krippendorff, Schachner — 2023

...

This talk ::

- Construct **all** flux vacua with $N_{\max} \leq 10$ for a simple example.



1. motivation
2. flux vacua
3. the mirror octic
4. results
5. summary

1. motivation
2. flux vacua
3. the mirror octic
4. results
5. summary

flux vacua

The setting.

Consider **type IIB** orientifold compactifications on **Calabi-Yau three-folds** $\mathcal{X}$ with O3-/O7-planes.

The four dimensional **effective theory** contains massless scalar fields (moduli) ::

- 1 axio-dilaton $\tau = c + i s ,$
- $h_-^{2,1}$ complex-structure moduli $z^i ,$
- ...

Three-form **fluxes** H and F along the compact space generate a potential for the moduli ::

$$W = \int_{\mathcal{X}} \Omega \wedge G , \quad \Omega \in H_-^{3,0}(\mathcal{X}) ,$$
$$G = F - \tau H .$$

Fluxes are constrained by the **tadpole cancellation condition**

$$0 = N_{\text{flux}} + 2N_{\text{D3}} + Q_{\text{D3}}.$$

The diagram illustrates the components of the tadpole cancellation condition equation $0 = N_{\text{flux}} + 2N_{\text{D3}} + Q_{\text{D3}}$. Arrows from the terms in the equation point to their respective definitions:

- N_{flux} points to a box containing "flux number $N_{\text{flux}} = \int_{\mathcal{X}} F \wedge H$ ".
- $2N_{\text{D3}}$ points to a box containing "number of D3-branes".
- Q_{D3} points to a box containing "contribution of O3-/O7-planes & D7-branes — typically negative".

Minima of the scalar potential are determined by **vanishing F-terms**

$$\begin{aligned} F_{\tau} &= 0 \\ F_{zi} &= 0 \end{aligned} \quad \Longleftrightarrow \quad G = -i \star G.$$

flux vacua

Some relations.

For three-forms on $\mathcal{X}$ one can introduce an integral **symplectic basis** as (with $I = 0, \dots, h_-^{2,1}$)

$$\{\alpha_I, \beta^I\} \in H_-^3(\mathcal{X}, \mathbb{Z}) ,$$

$$\eta = \int_{\mathcal{X}} \begin{pmatrix} \alpha \\ \beta \end{pmatrix} \wedge (\alpha, \beta) = \begin{pmatrix} 0 & +\mathbb{1} \\ -\mathbb{1} & 0 \end{pmatrix} ,$$

$$\mathcal{M} = \int_{\mathcal{X}} \begin{pmatrix} \alpha \\ \beta \end{pmatrix} \wedge \star(\alpha, \beta) .$$

The matrix $\mathcal{M}$ is symplectic, symmetric, and positive-definite. Its **eigenvalues** come in pairs

$$(\lambda_I, \lambda_I^{-1}) , \qquad \lambda_I \geq 1 .$$

The **fluxes** can be expanded in the symplectic basis and combined into vectors as

$$H = h^I \alpha_I + h_I \beta^I \qquad \mathbf{H} = \begin{pmatrix} h^I \\ h_I \end{pmatrix} ,$$

$$F = f^I \alpha_I + f_I \beta^I \qquad \mathbf{F} = \begin{pmatrix} f^I \\ f_I \end{pmatrix} .$$

The **minimum condition** can be expressed in the following way

$$G = -i \star G \quad \longrightarrow \quad \eta(F - Hc) = -\mathcal{M}Hs.$$

Using this condition one obtains the following relations

$$s = \frac{N_{\text{flux}}}{H^T \mathcal{M} H}, \quad c = \frac{H^T \mathcal{M} F}{H^T \mathcal{M} H}, \quad F^T \mathcal{M} F = (s^2 + c^2)(H^T \mathcal{M} H).$$

The usual **bounds** on matrix norms imply ($\|\cdot\|^2$ denotes the Euclidean norm)

$$\frac{1}{\lambda_{\max}} \|\mathbf{H}\|^2 \leq H^T \mathcal{M} H \leq \lambda_{\max} \|\mathbf{H}\|^2.$$

flux vacua

Finite fluxes in finite regions.

Using **S-duality**, the axio-dilaton can be mapped into the fundamental domain. This implies

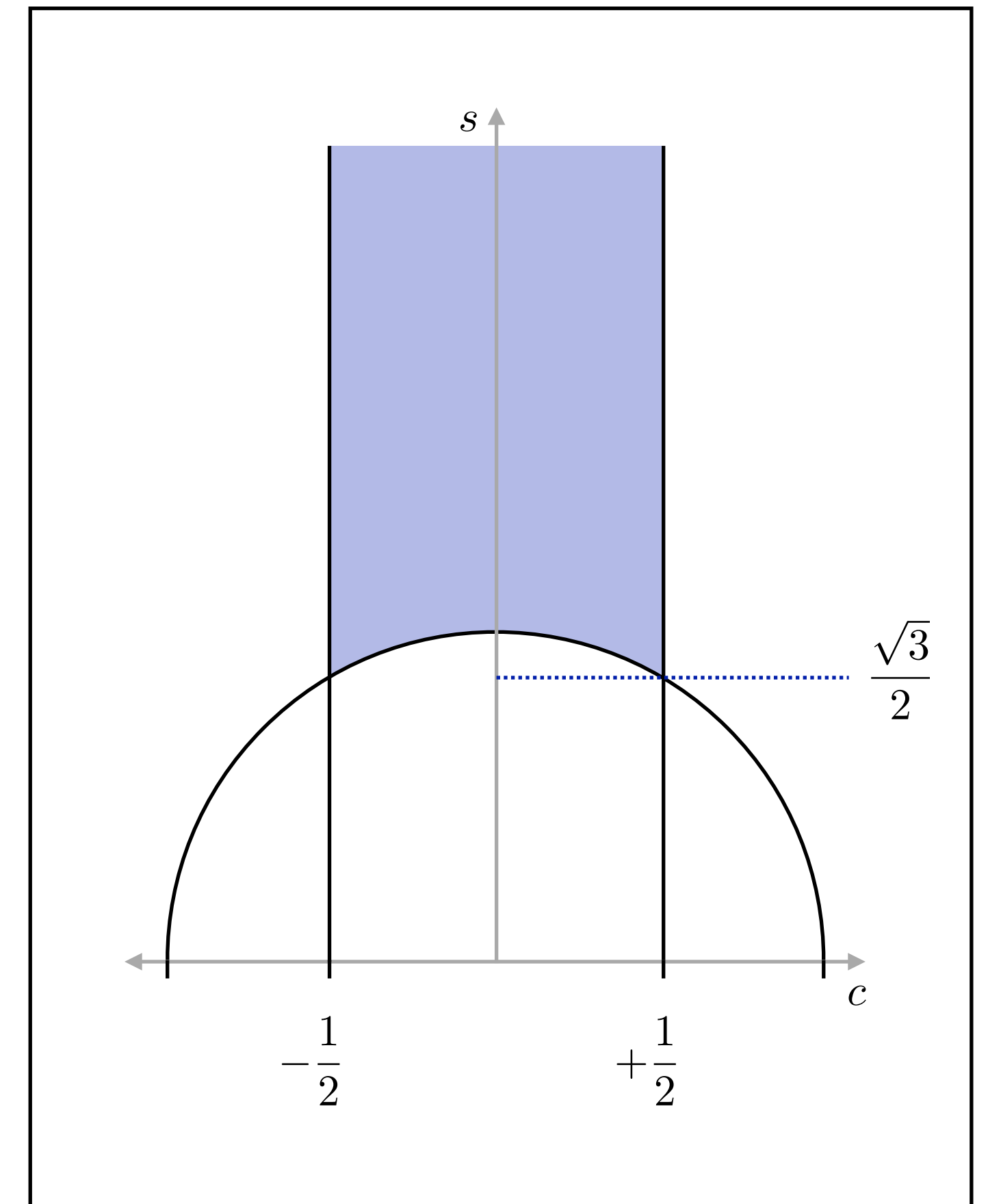
$$\frac{\sqrt{3}}{2} \leq s$$

$$\frac{\sqrt{3}}{2} \leq \frac{N_{\text{flux}}}{\mathbf{H}^T \mathcal{M} \mathbf{H}}$$

$$\frac{\sqrt{3}}{2} \leq \frac{N_{\text{flux}}}{\mathbf{H}^T \mathcal{M} \mathbf{H}} \leq N_{\text{flux}} \frac{\lambda_{\text{max}}}{\|\mathbf{H}\|^2}$$



$$\|\mathbf{H}\|^2 \leq \frac{2N_{\text{flux}} \lambda_{\text{max}}}{\sqrt{3}}.$$



With the axio-dilaton in the **fundamental domain** one also obtains

$$(\mathbf{F}^T \mathcal{M} \mathbf{F}) = (c^2 + s^2)(\mathbf{H}^T \mathcal{M} \mathbf{H})$$

$$(\mathbf{H}^T \mathcal{M} \mathbf{H})(\mathbf{F}^T \mathcal{M} \mathbf{F}) = (c^2 + s^2)(\mathbf{H}^T \mathcal{M} \mathbf{H})^2$$

$$(\mathbf{H}^T \mathcal{M} \mathbf{H})(\mathbf{F}^T \mathcal{M} \mathbf{F}) = (c^2 + s^2) \frac{N_{\text{flux}}^2}{s^2}$$

$$\frac{\|\mathbf{H}\|^2}{\lambda_{\text{max}}} \frac{\|\mathbf{F}\|^2}{\lambda_{\text{max}}} \leq (\mathbf{H}^T \mathcal{M} \mathbf{H})(\mathbf{F}^T \mathcal{M} \mathbf{F}) = (c^2 + s^2) \frac{N_{\text{flux}}^2}{s^2} \leq \frac{4}{3} N_{\text{flux}}^2$$



$$\|\mathbf{F}\|^2 \leq \frac{4N_{\text{flux}}^2 \lambda_{\text{max}}^2}{3\|\mathbf{H}\|^2}.$$

Summary ::

- For a region of complex-structure moduli space in which the eigenvalues of the Hodge-star operator are bounded,
- and for a given flux number N_{flux} ,
- only **finitely-many** flux choices can lead to a vacuum in that region.

Remark ::

- The bounds can be made slightly stronger.

1. motivation
2. flux vacua
3. the mirror octic
4. results
5. summary

the mirror octic

The periods.

The **octic three-fold** is a hyper-surface in weighted projective space $\mathbb{P}_{11114}$ defined by

$$\sum_{i=1}^4 x_i^8 + 4x_0^2 - 8\psi x_0 x_1 x_2 x_3 x_4 = 0.$$

The **mirror octic** has Hodge numbers $(h^{1,1}, h^{2,1}) = (149, 1)$.

For the mirror octic an **orientifold projection** can be chosen such that

$$\begin{array}{llll} h_-^{2,1} = 1, & h_+^{2,1} = 0, & N_{O7} = 0, & \\ h_-^{1,1} = 72, & h_+^{1,1} = 77, & N_{O3} = 16, & Q_{D3} = -8. \end{array}$$

The **octic three-fold** is a hyper-surface in weighted projective space $\mathbb{P}_{11114}$ defined by

$$\sum_{i=1}^4 x_i^8 + 4x_0^2 - 8\psi x_0 x_1 x_2 x_3 x_4 = 0.$$

The **mirror octic** has Hodge numbers $(h^{1,1}, h^{2,1}) = (149, 1)$.

For the mirror octic an **orientifold projection** can be chosen such that

$$h_-^{2,1} = 1,$$

$$h_-^{1,1} = 72,$$

$$h_+^{2,1} = 0,$$

$$h_+^{1,1} = 77,$$

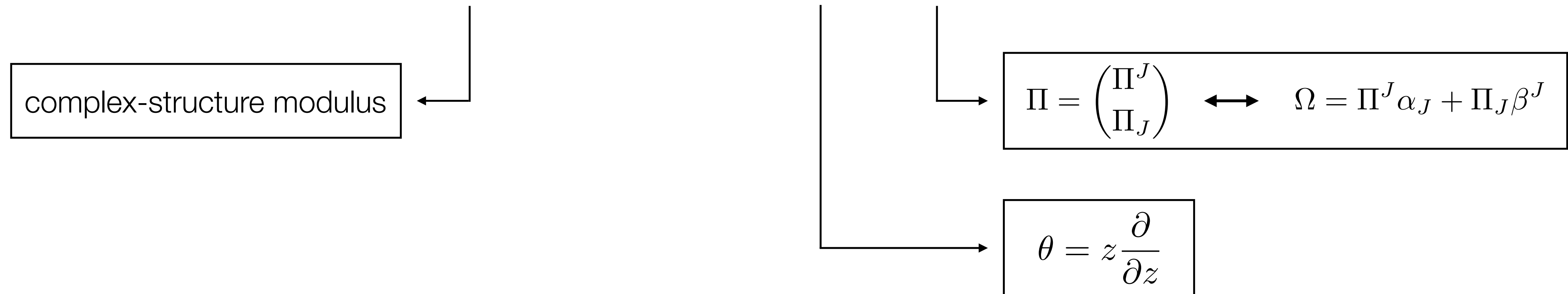
$$N_{O7} = 0,$$

$$N_{O3} = 16,$$

$$Q_{D3} = -8.$$

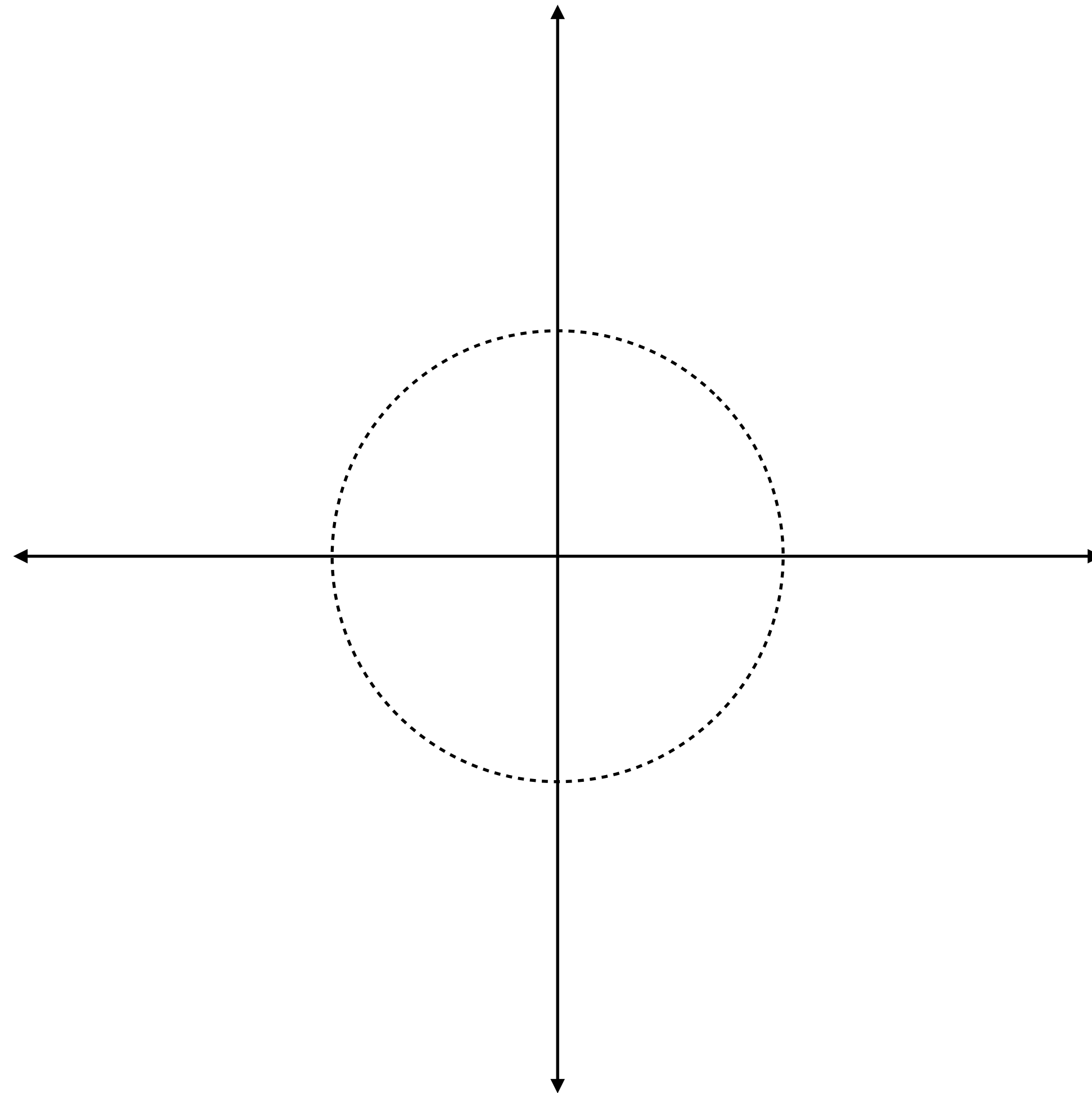
The periods that determine the holomorphic three-form are solutions to the **Picard-Fuchs equation**

$$\left[\theta^4 - z \left(\theta + \frac{1}{8} \right) \left(\theta + \frac{3}{8} \right) \left(\theta + \frac{5}{8} \right) \left(\theta + \frac{7}{8} \right) \right] \Pi = 0.$$

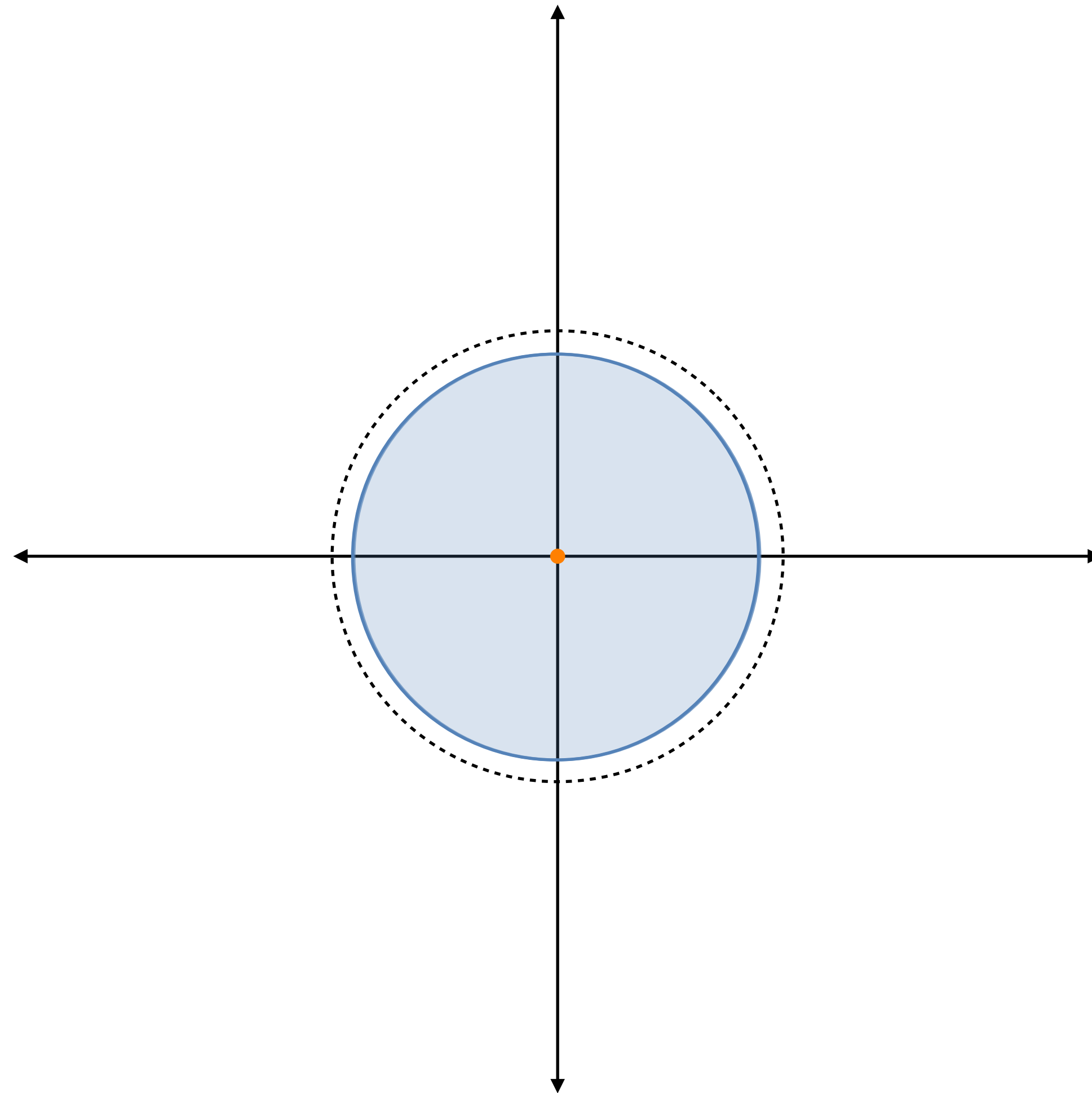


The solutions are typically expanded in power series that converge for $|z| < 1$.

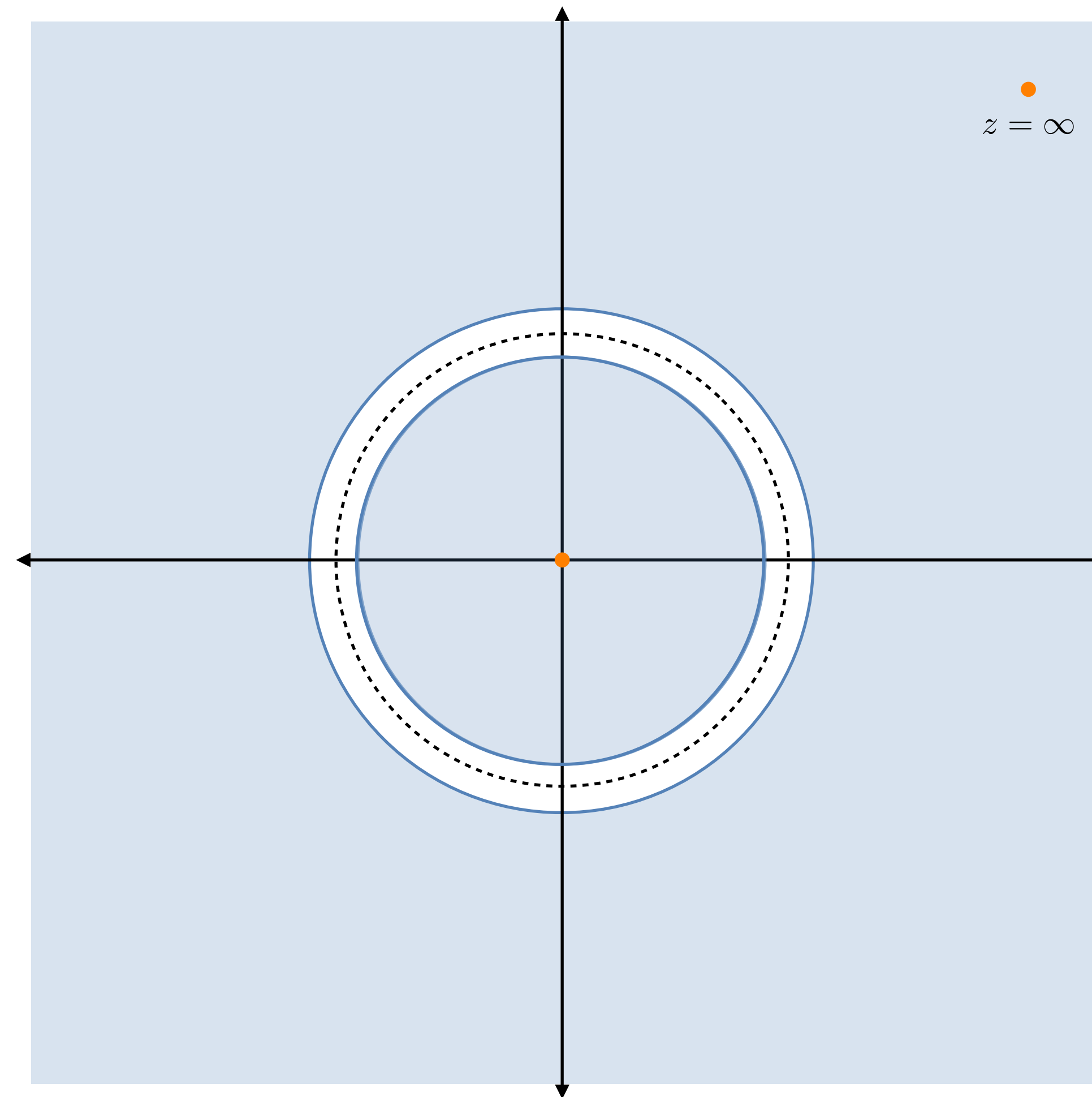
The complex-structure moduli space can be covered by **six discs** of radius 0.9.



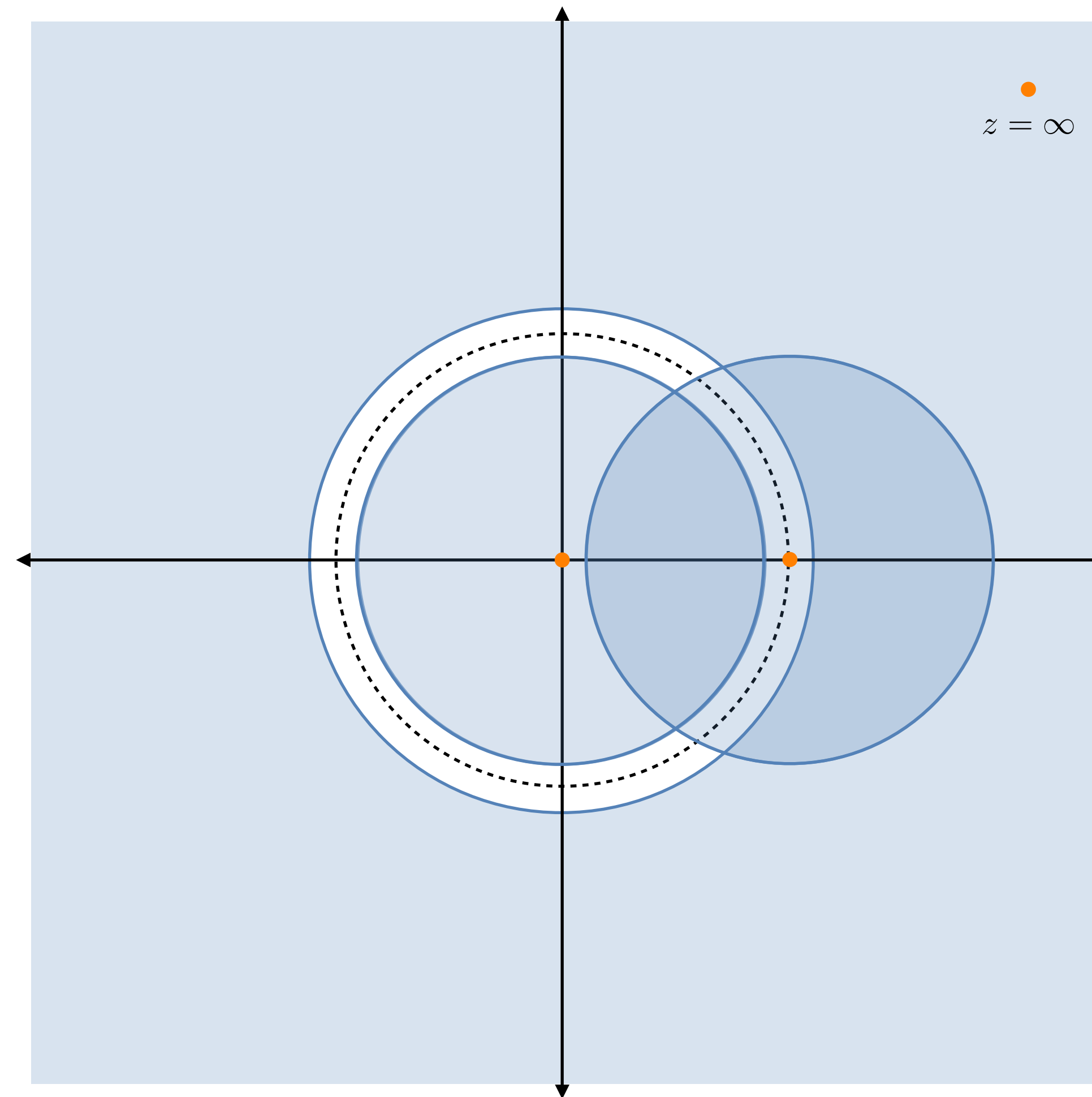
The complex-structure moduli space can be covered by **six discs** of radius 0.9.



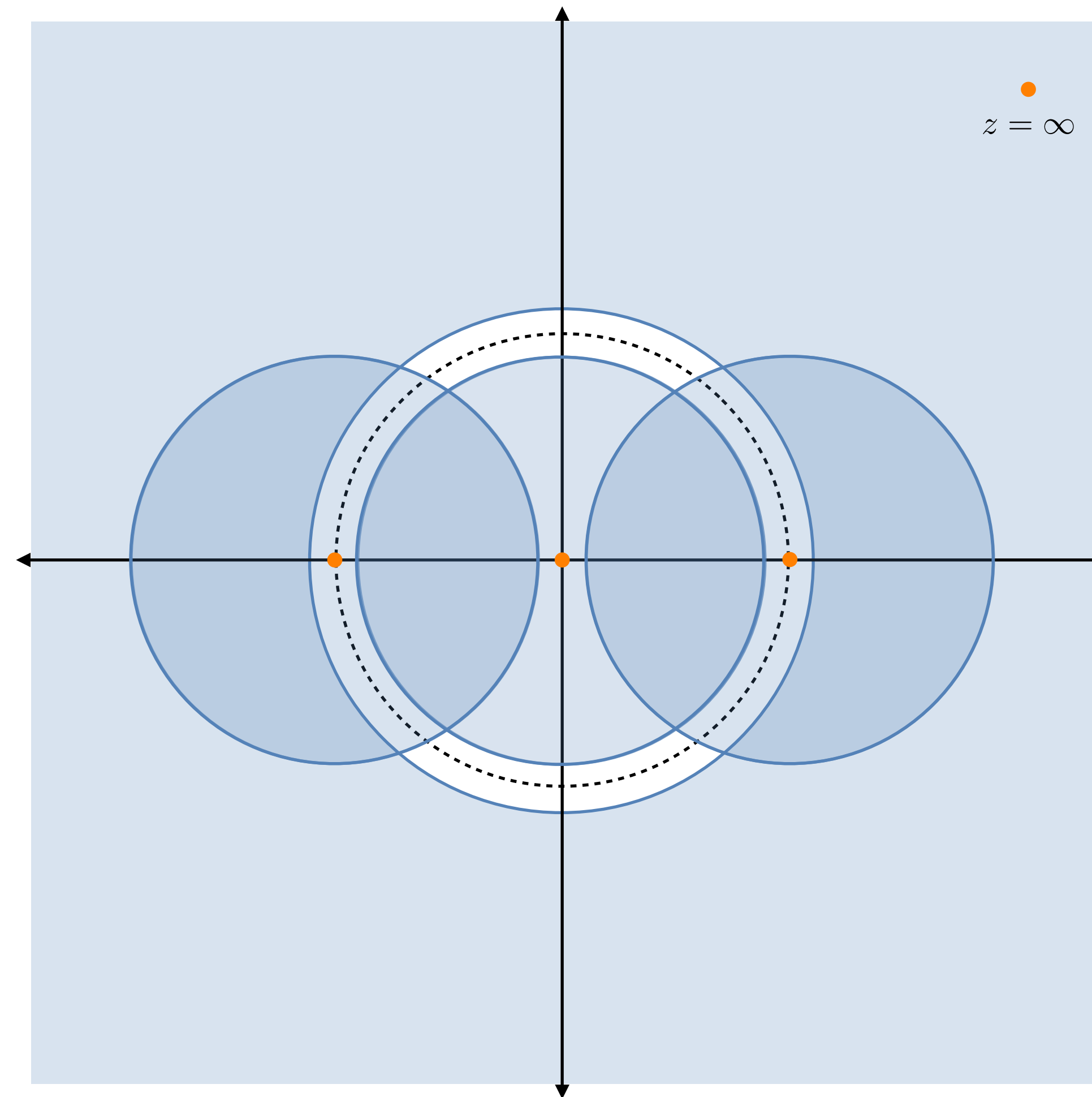
The complex-structure moduli space can be covered by **six discs** of radius 0.9.



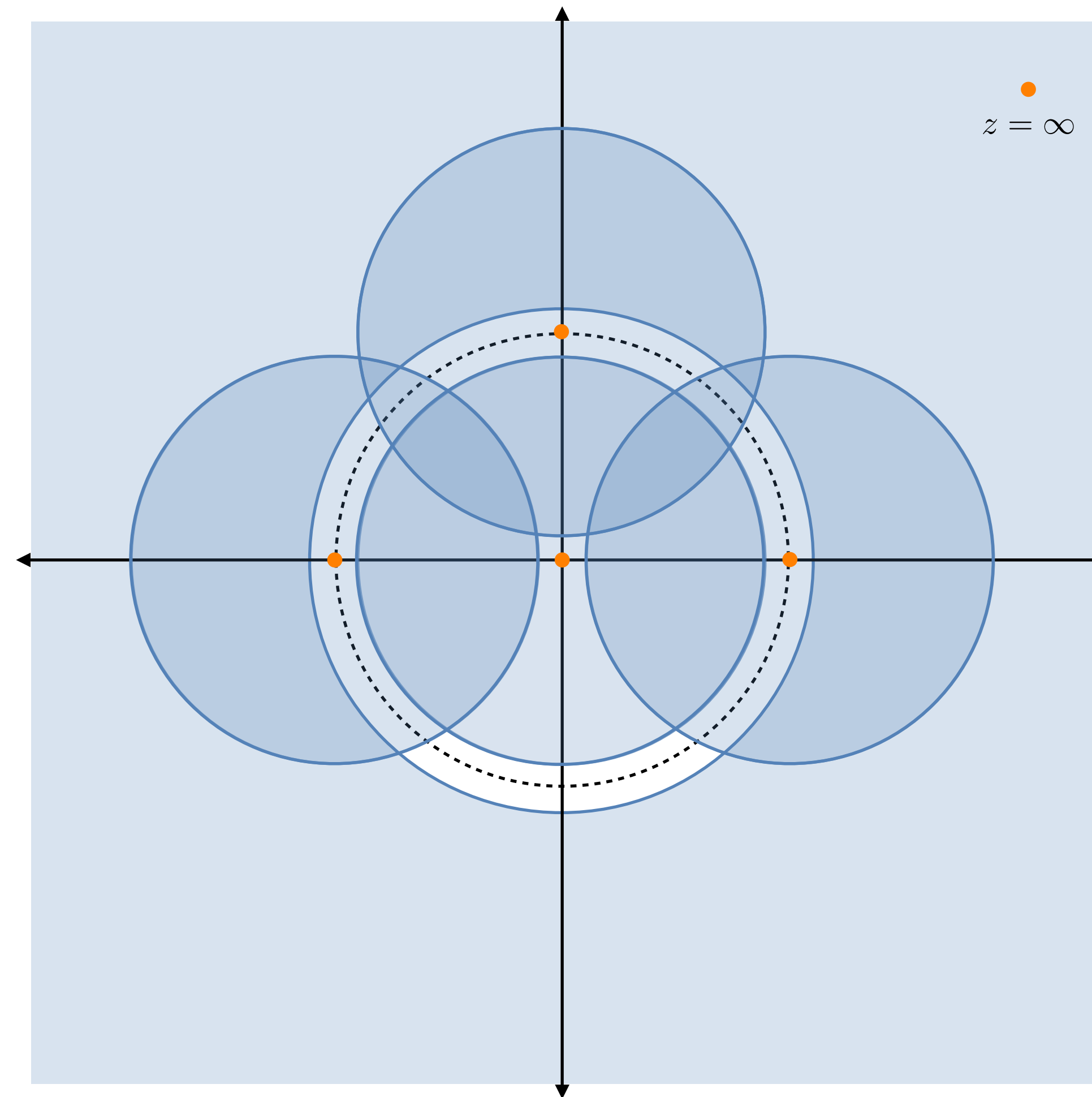
The complex-structure moduli space can be covered by **six discs** of radius 0.9.



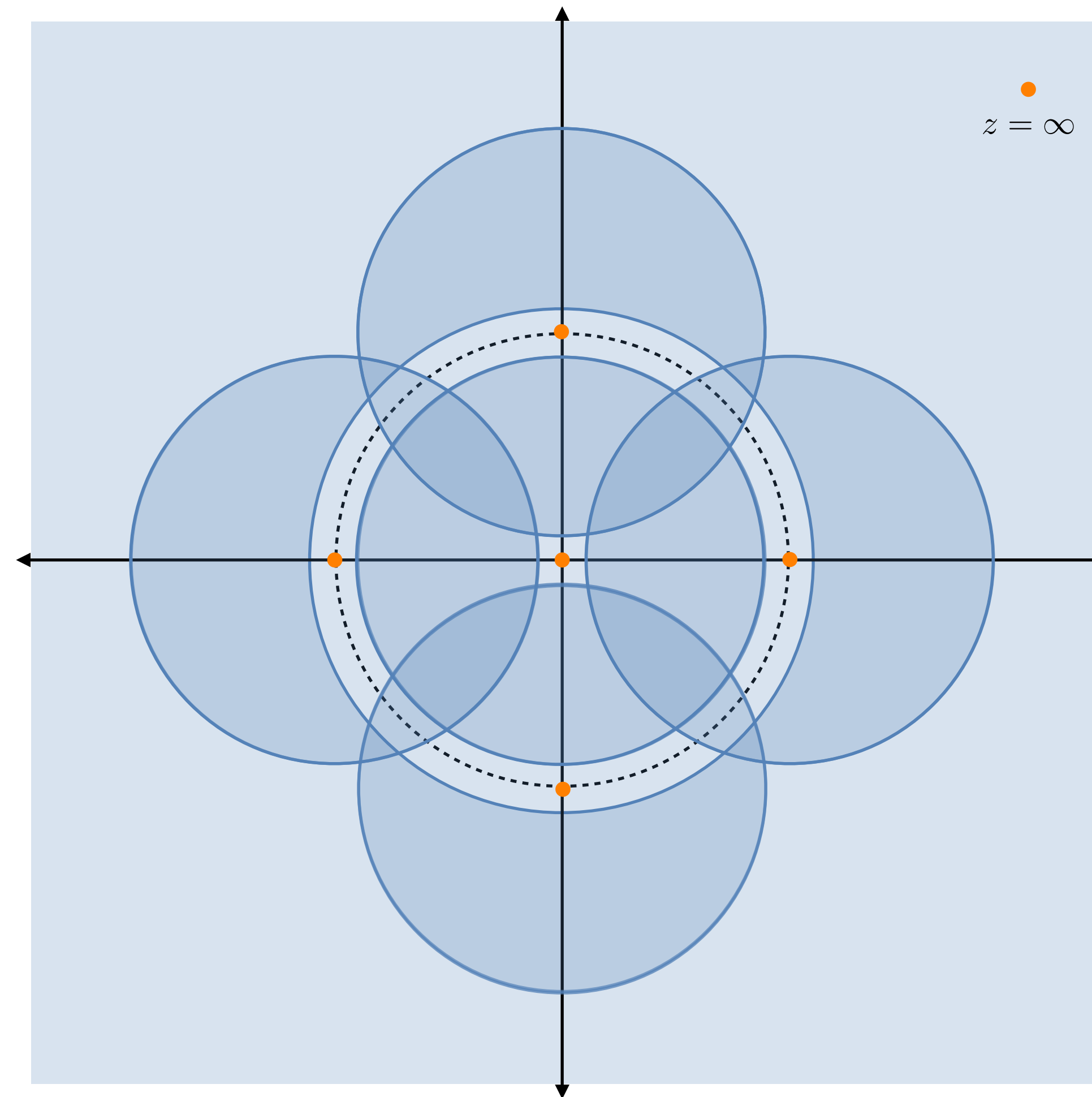
The complex-structure moduli space can be covered by **six discs** of radius 0.9.



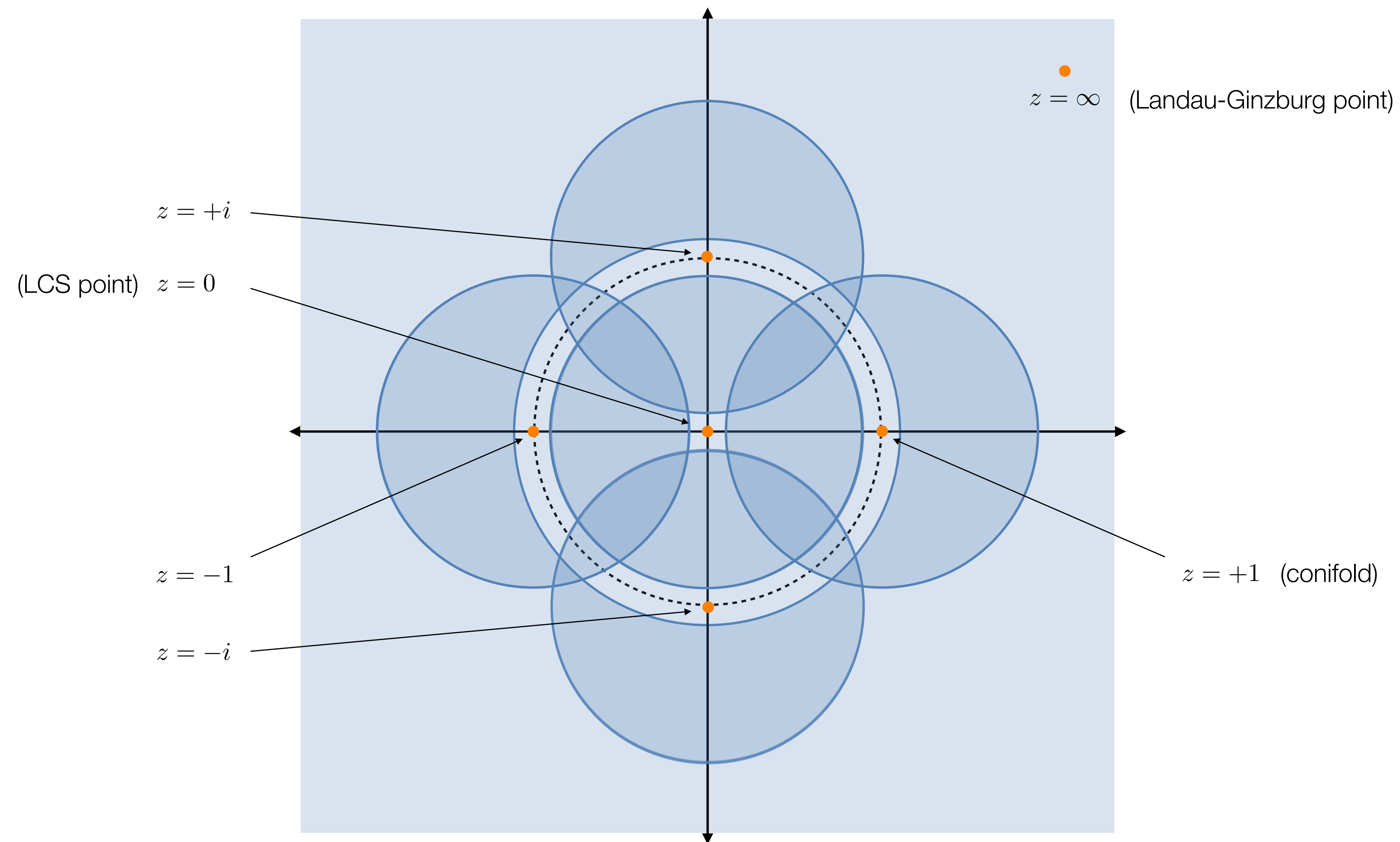
The complex-structure moduli space can be covered by **six discs** of radius 0.9.



The complex-structure moduli space can be covered by **six discs** of radius 0.9.

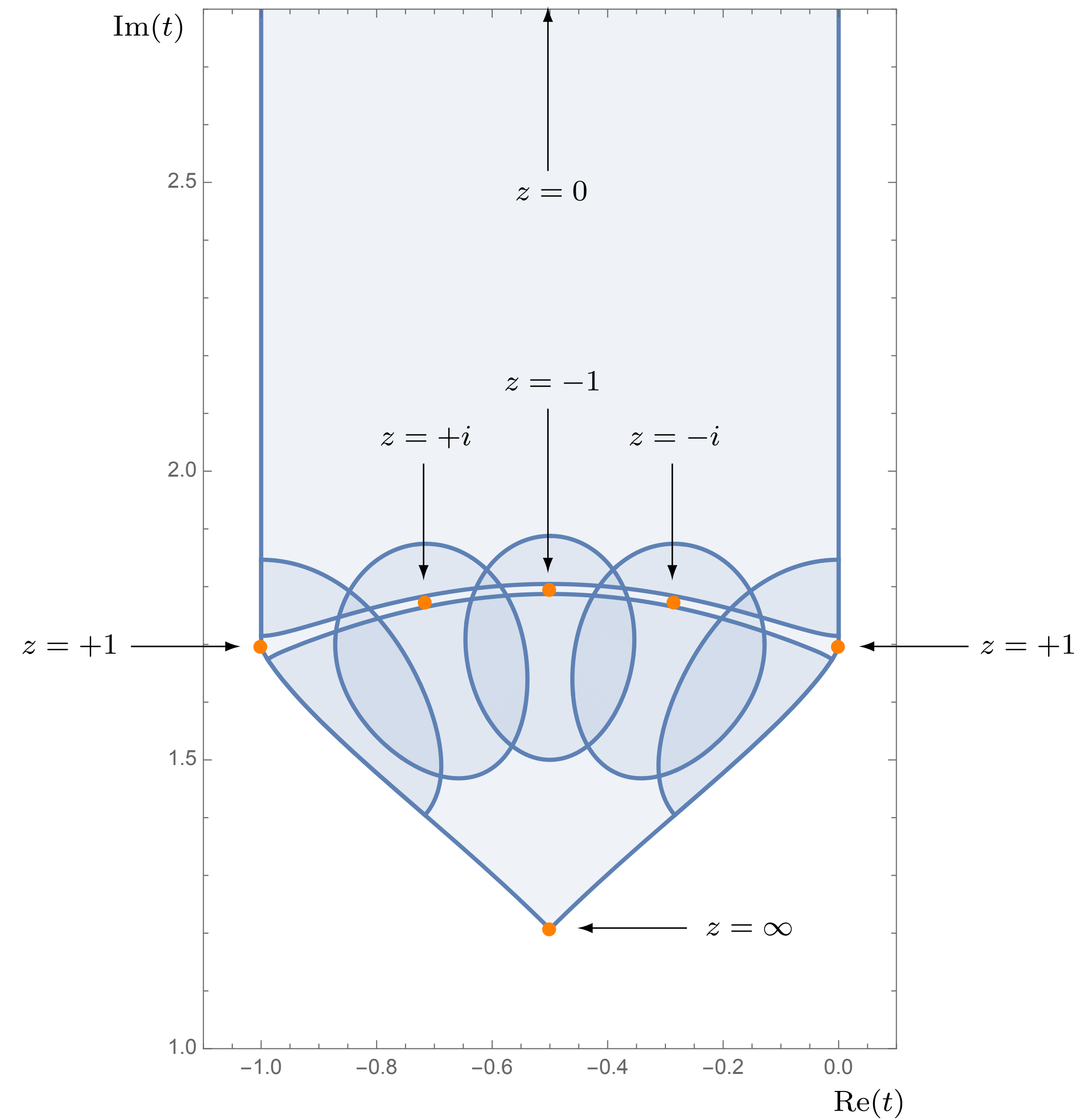


The complex-structure moduli space can be covered by **six discs** of radius 0.9.



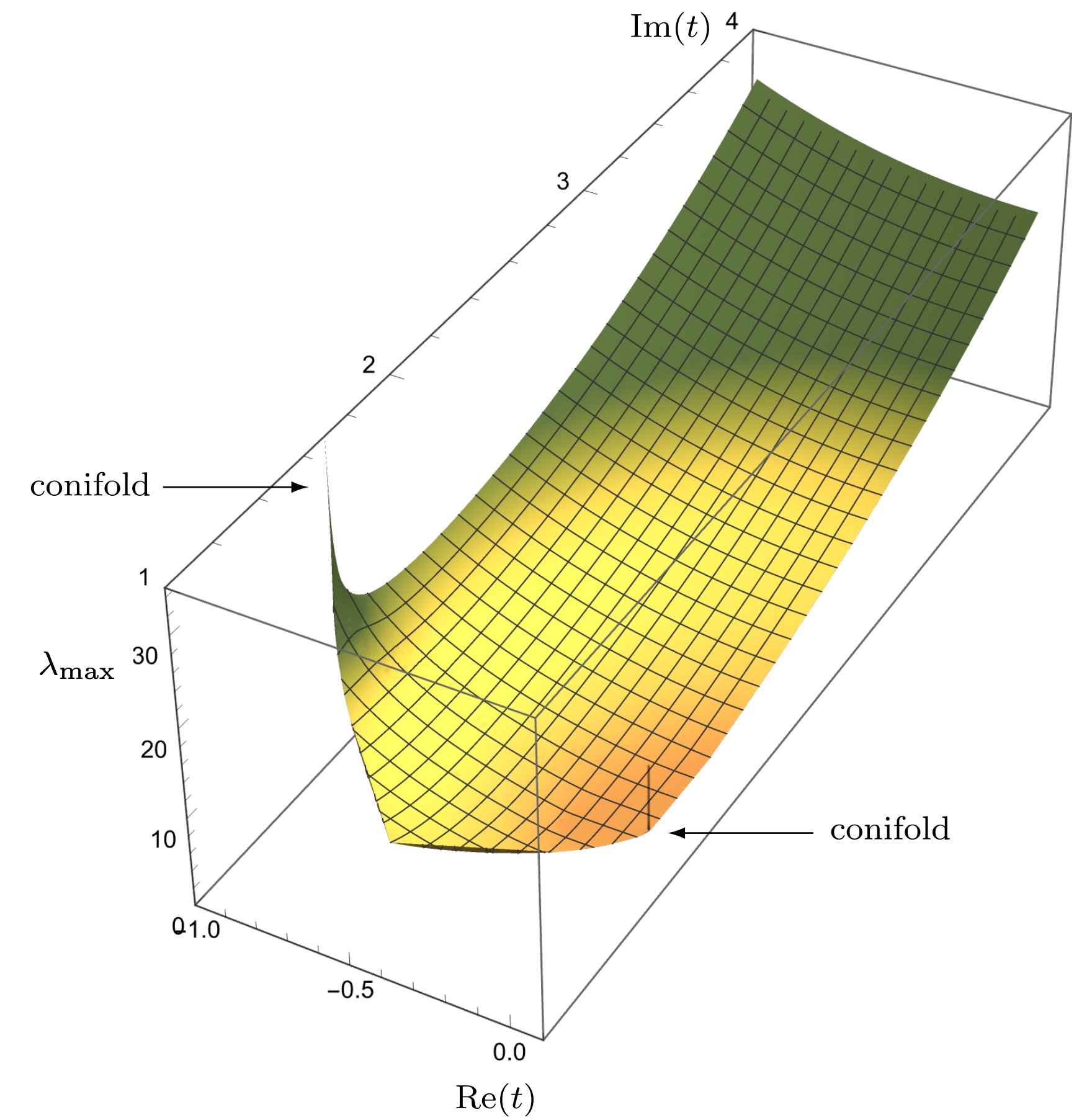
For convenience, the moduli space can be mapped to the Kähler side via the **mirror map**

$$t(z) = \frac{\Pi^2(z)}{\Pi^1(z)}.$$



Strategy flux vacua.

The **largest eigenvalue** $\lambda_{\max}$ of the Hodge-star operator diverges at the conifold and LCS point.



For each **bulk region** ($\lambda_{\max}$ is finite) ::

- Determine periods Π up to $\mathcal{O}(100)$.
- Determine relevant flux choices for $N_{\max}$.
- Try to solve minimum condition.

point	region	$\lambda_{\max}$
$z_{\text{ep}} = 0$	$10^{-9} \leq z - z_{\text{ep}} \leq 0.9$	37.98
$z_{\text{ep}} = +1$	$10^{-9} \leq z - z_{\text{ep}} \leq 0.9$	3.95
$z_{\text{ep}} = \infty$	$ z - z_{\text{ep}} \leq 0.9$	8.33
$z_{\text{ep}} = -1$	$ z - z_{\text{ep}} \leq 0.9$	5.00
$z_{\text{ep}} = +i$	$ z - z_{\text{ep}} \leq 0.9$	11.24
$z_{\text{ep}} = -i$	$ z - z_{\text{ep}} \leq 0.9$	3.35

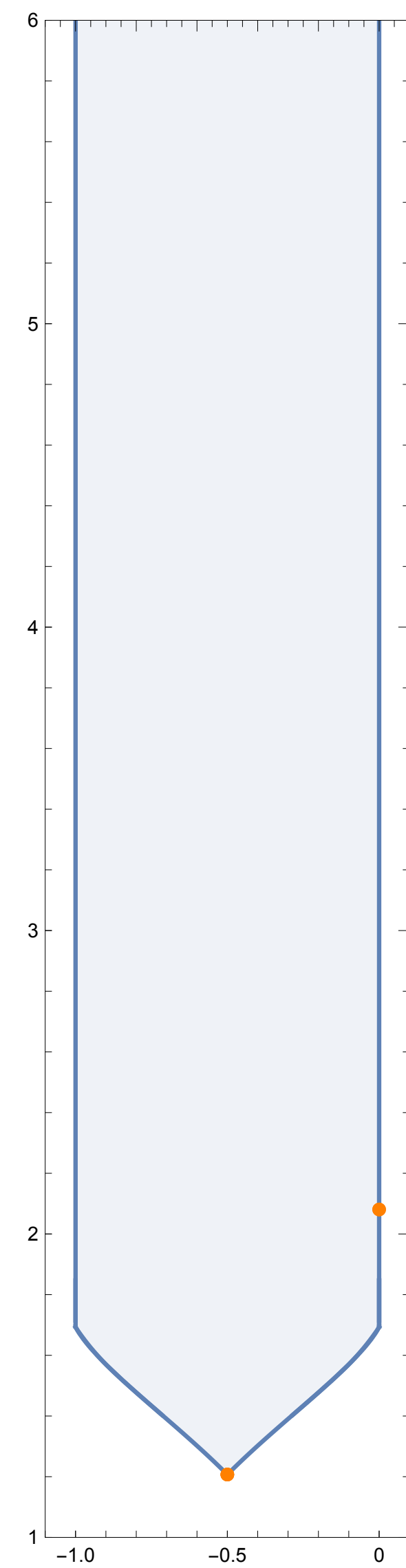
For each **boundary region** ($\lambda_{\max}$ diverges) ::

- Determine periods at leading order.
- Determine relevant flux choices for $N_{\max}$ (partially) analytically.
- Solve minimum condition.

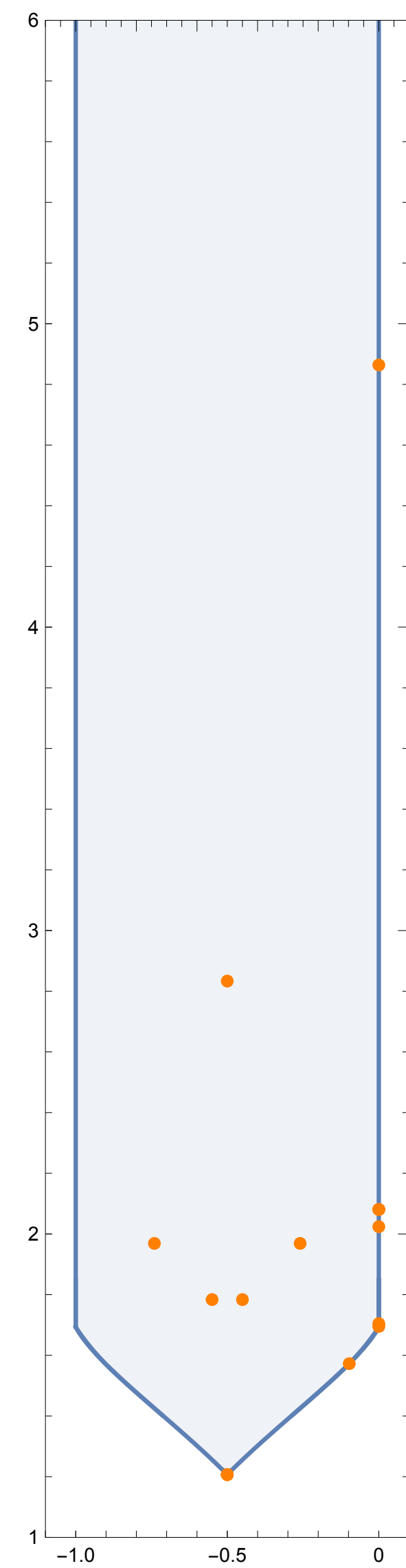
point	region
$z_{\text{ep}} = 0$	$ z - z_{\text{ep}} \leq 10^{-9}$
$z_{\text{ep}} = +1$	$ z - z_{\text{ep}} \leq 10^{-9}$

1. motivation
2. flux vacua
3. the mirror octic
4. results
5. summary

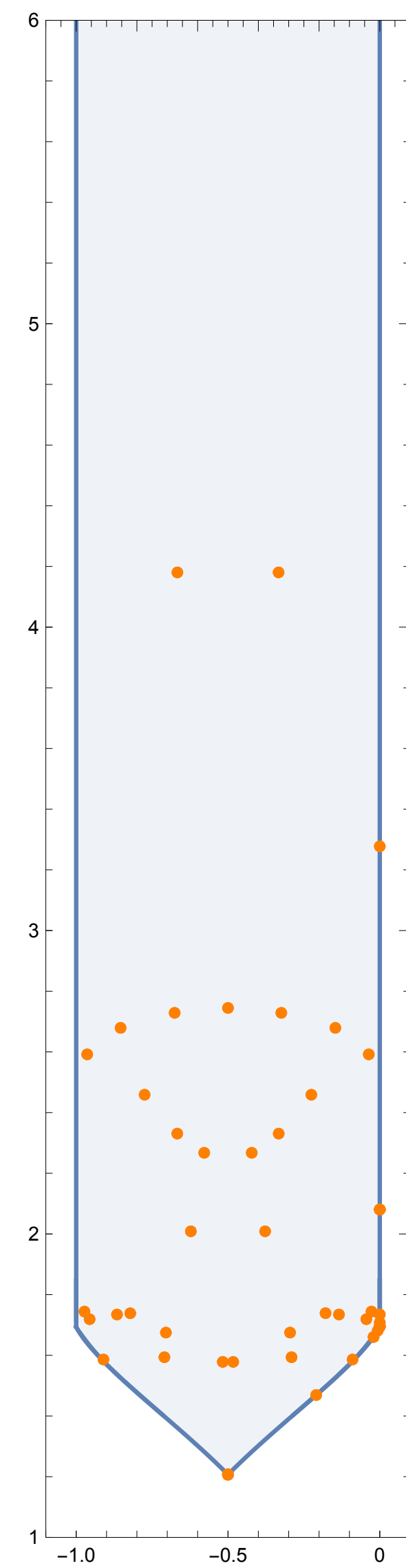
Distribution of vacua.



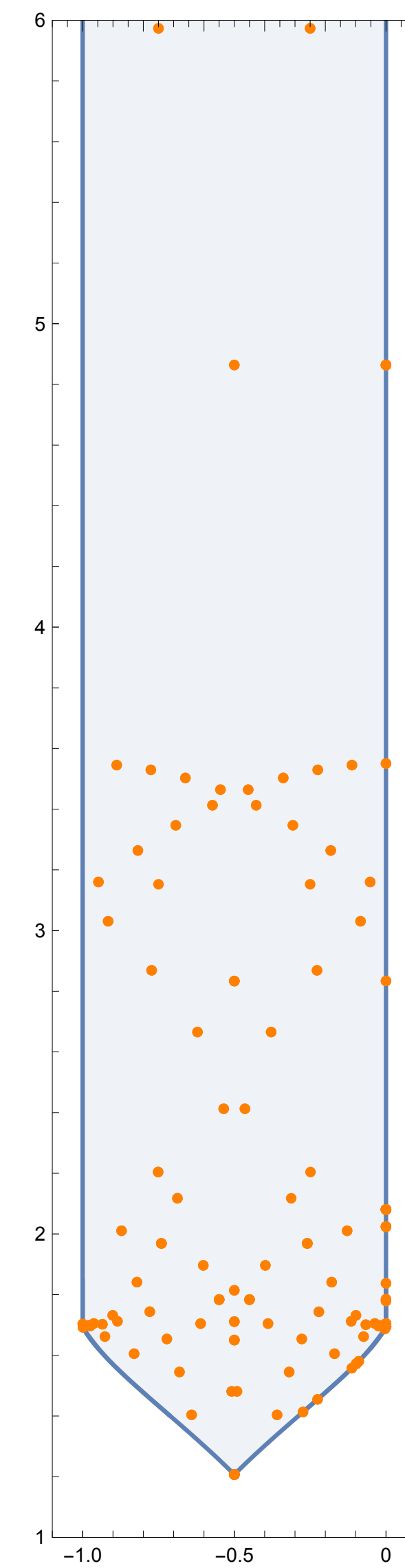
$N_{\text{flux}} = 1$



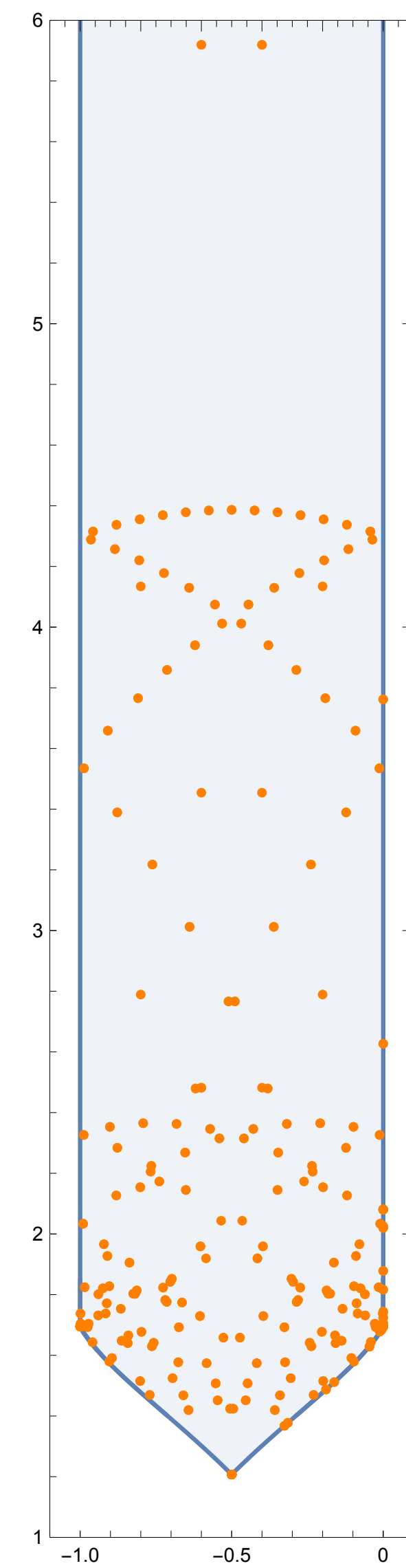
$N_{\text{flux}} = 2$



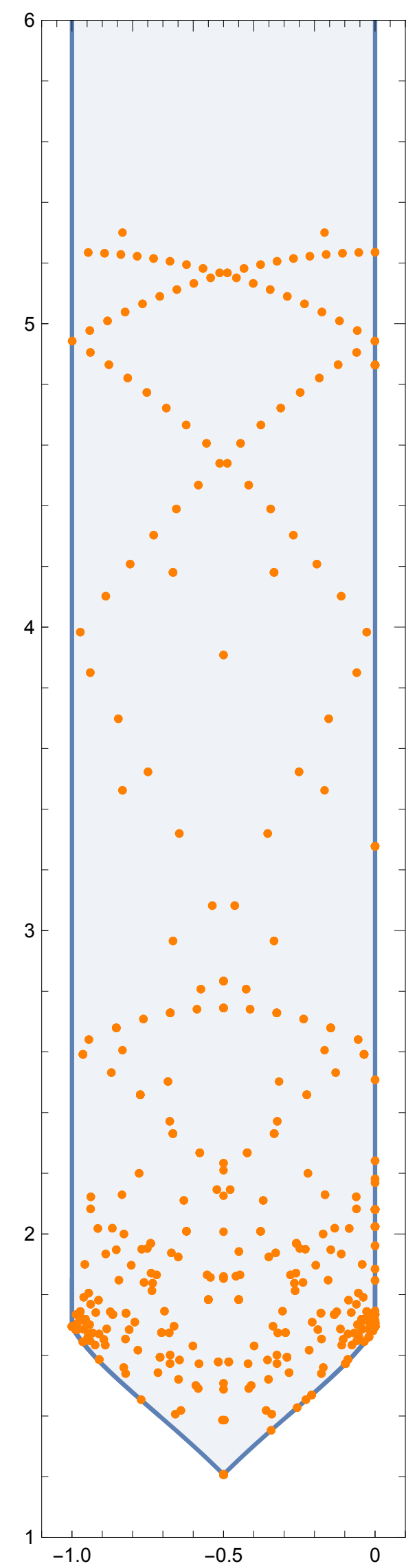
$N_{\text{flux}} = 3$



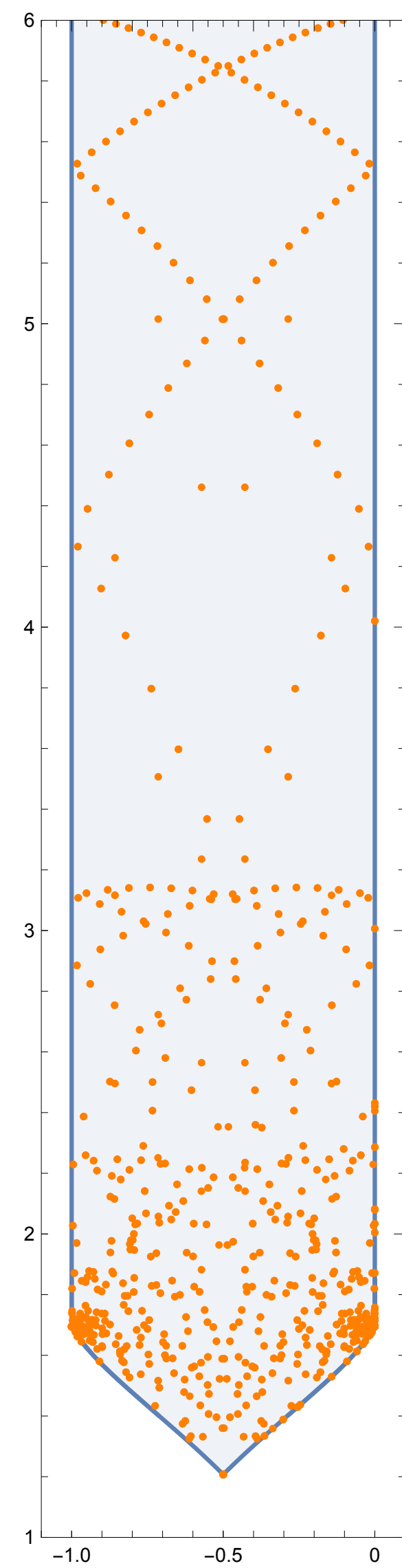
$N_{\text{flux}} = 4$



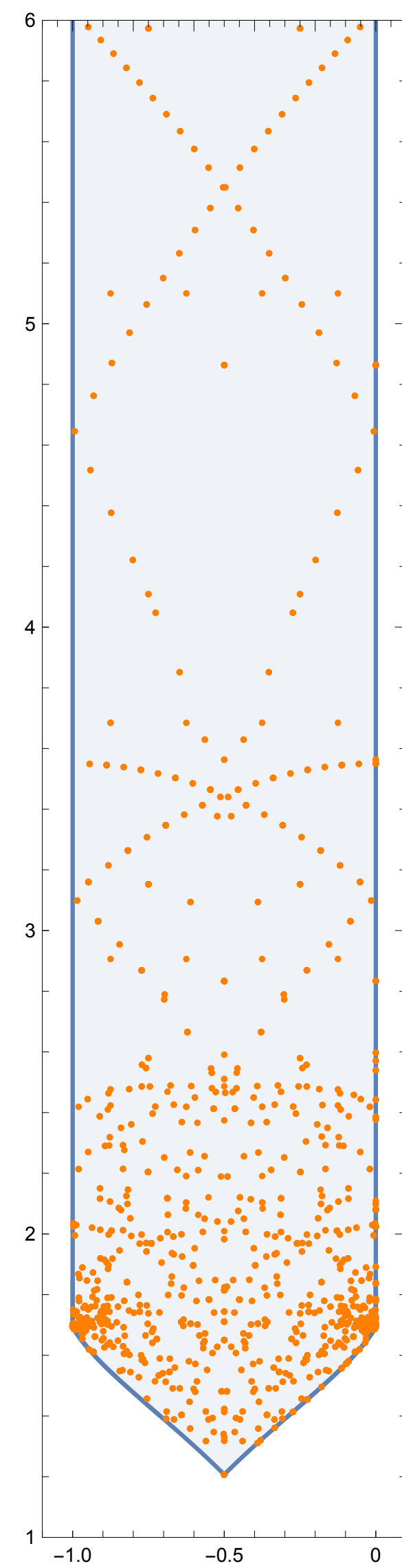
$N_{\text{flux}} = 5$



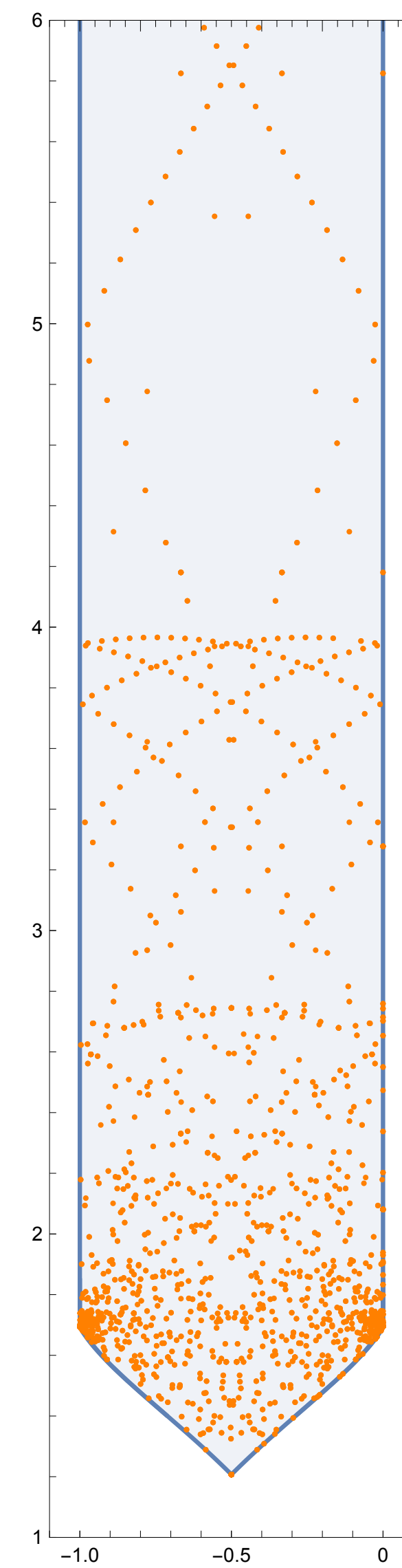
$N_{\text{flux}} = 6$



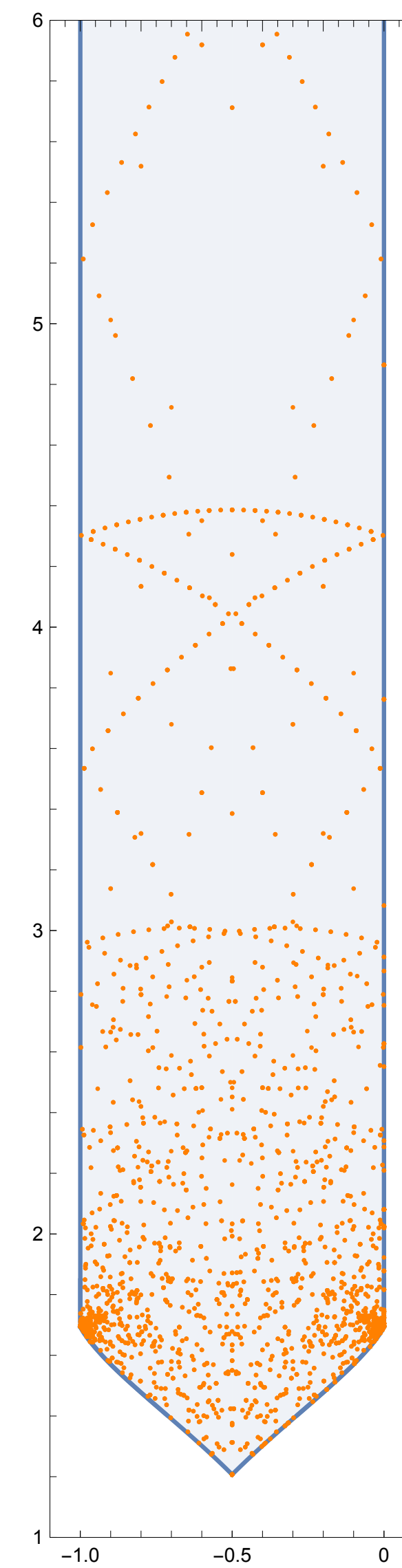
$N_{\text{flux}} = 7$



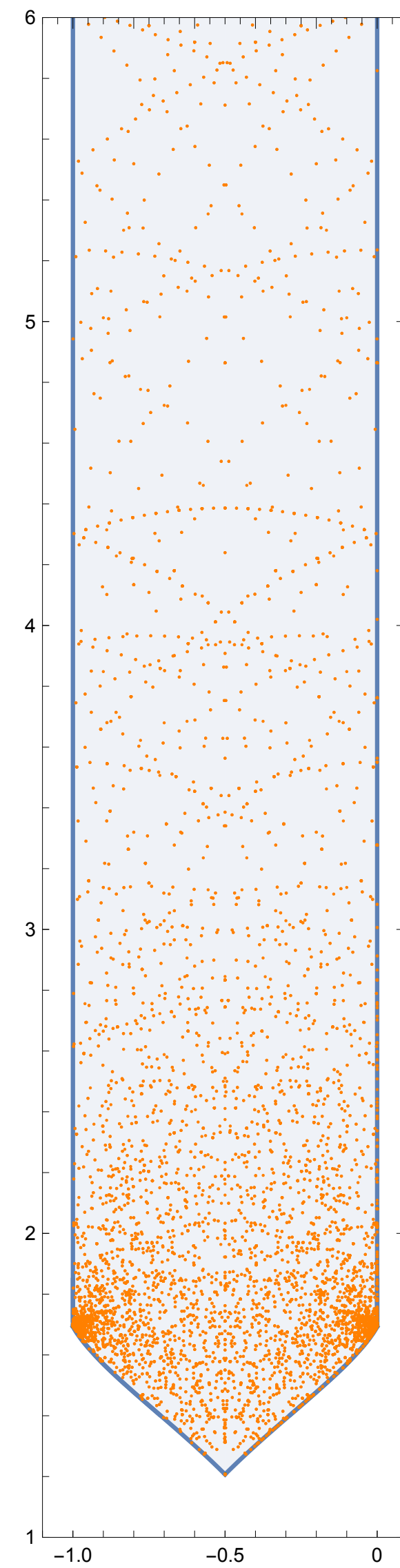
$N_{\text{flux}} = 8$



$N_{\text{flux}} = 9$



$N_{\text{flux}} = 10$



$$N_{\text{flux}} \leq 10$$

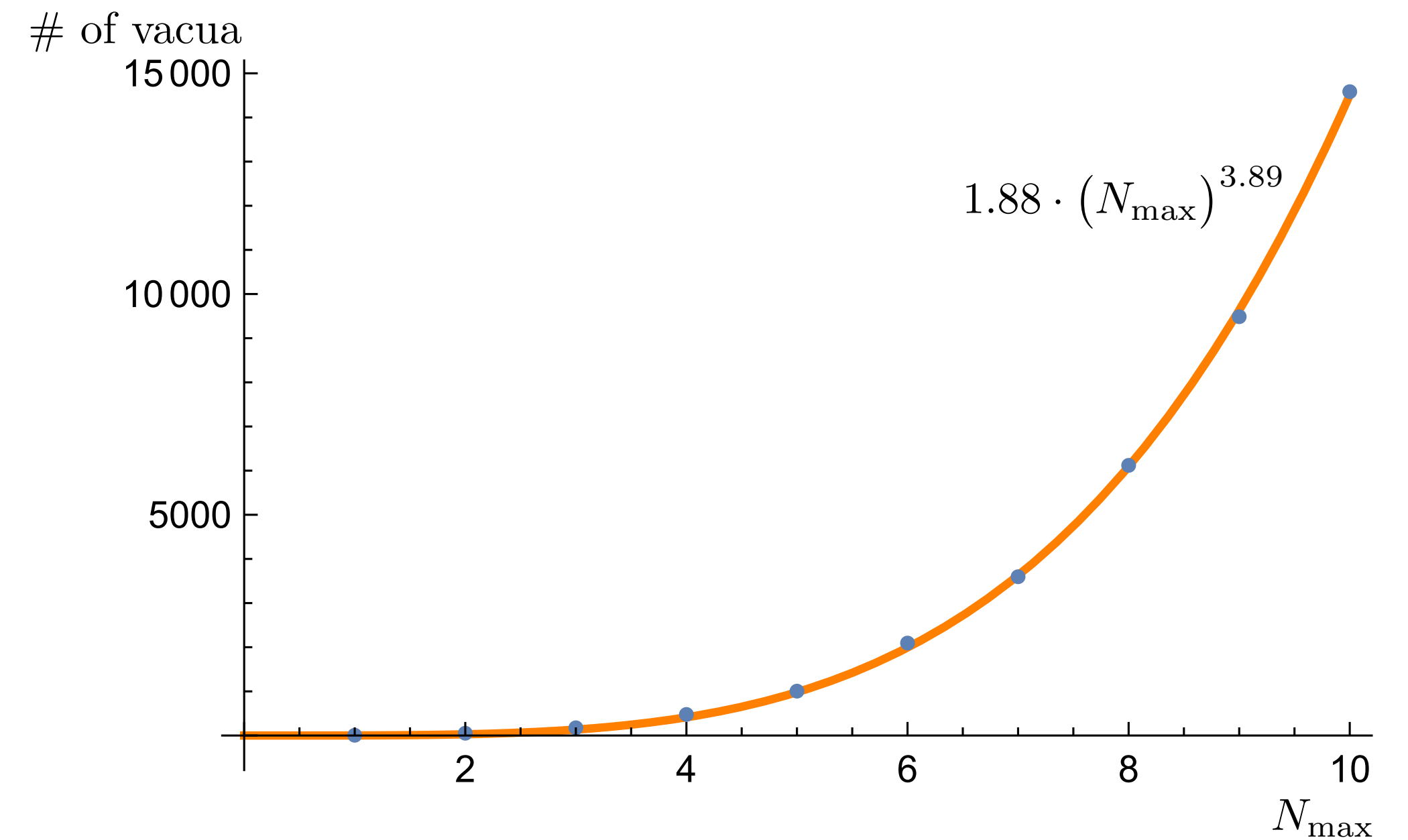
Number of vacua.

The cumulative **number of vacua** for a given $N_{\max}$ can be fitted as

$$\mathcal{N}(N_{\text{flux}} \leq N_{\max}) = 1.88 \cdot (N_{\max})^{3.89}.$$

This result agrees approximately with the estimate of Denef/Douglas

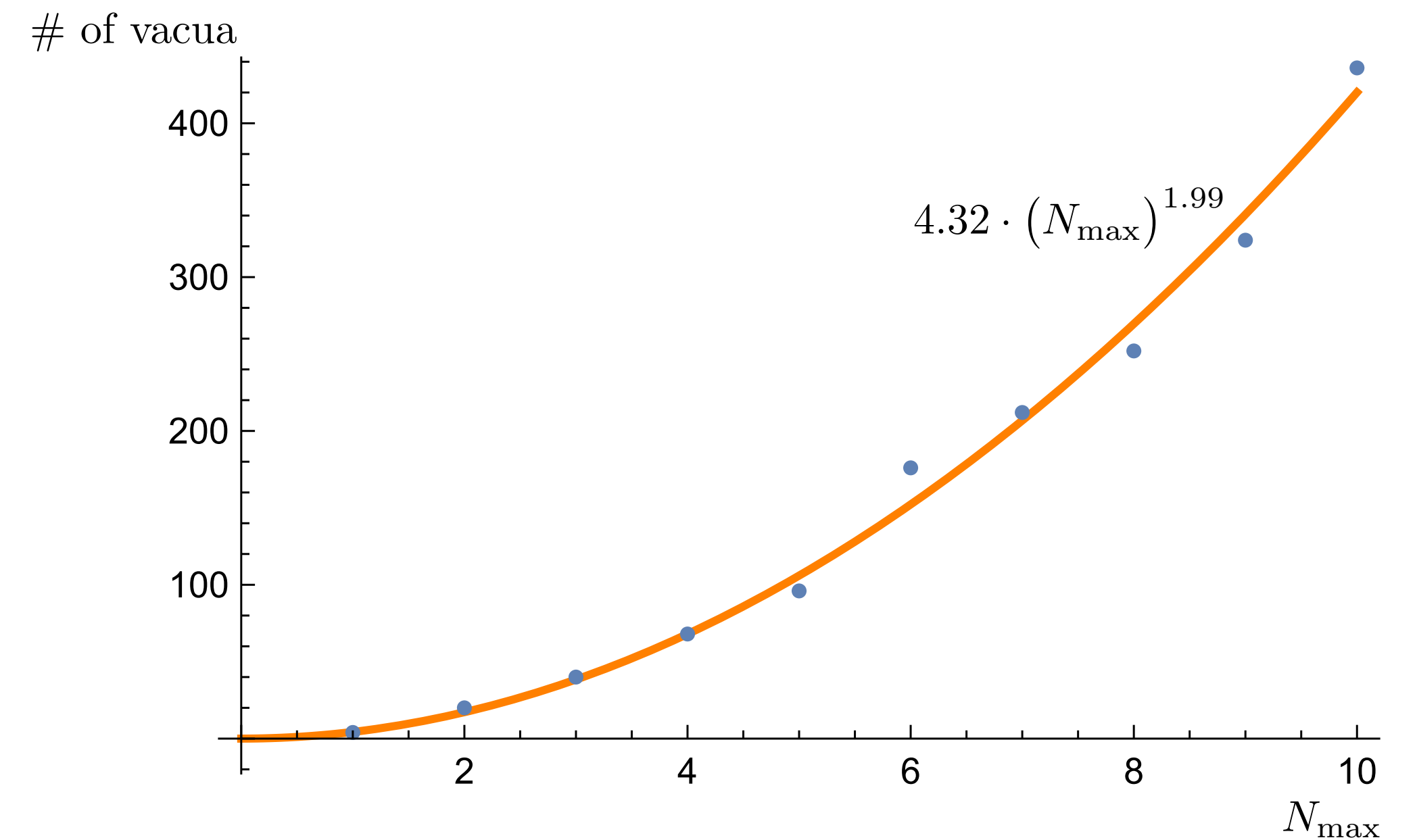
$$\mathcal{N}_{\text{est}}(N_{\text{flux}} \leq N_{\max}) = 1.37 \cdot (N_{\max})^4.$$



The cumulative **number of vacua** with $W_0 = 0$ for a given N_{max} can be fitted as

$$\mathcal{N}(N_{\text{flux}} \leq N_{\text{max}})_{W_0=0} = 4.32 \cdot (N_{\text{max}})^{1.99}.$$

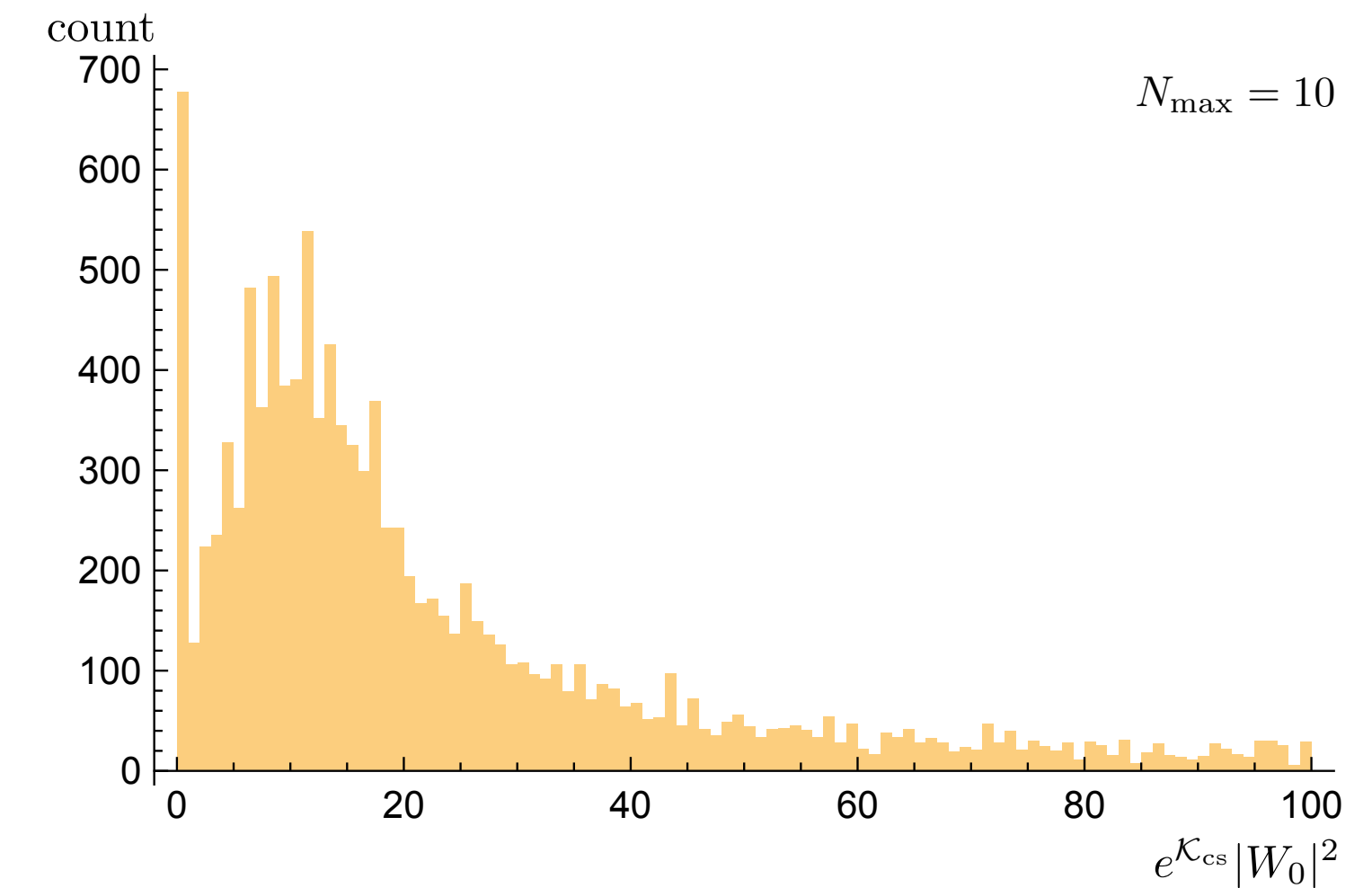
All these vacua are located at the **Landau-Ginzburg** point.



Distributions.

Relevant for moduli stabilization via [KKLT](#) and [LVS](#) is the quantity

$$e^{\mathcal{K}_{cs}} |W_0|^2 .$$

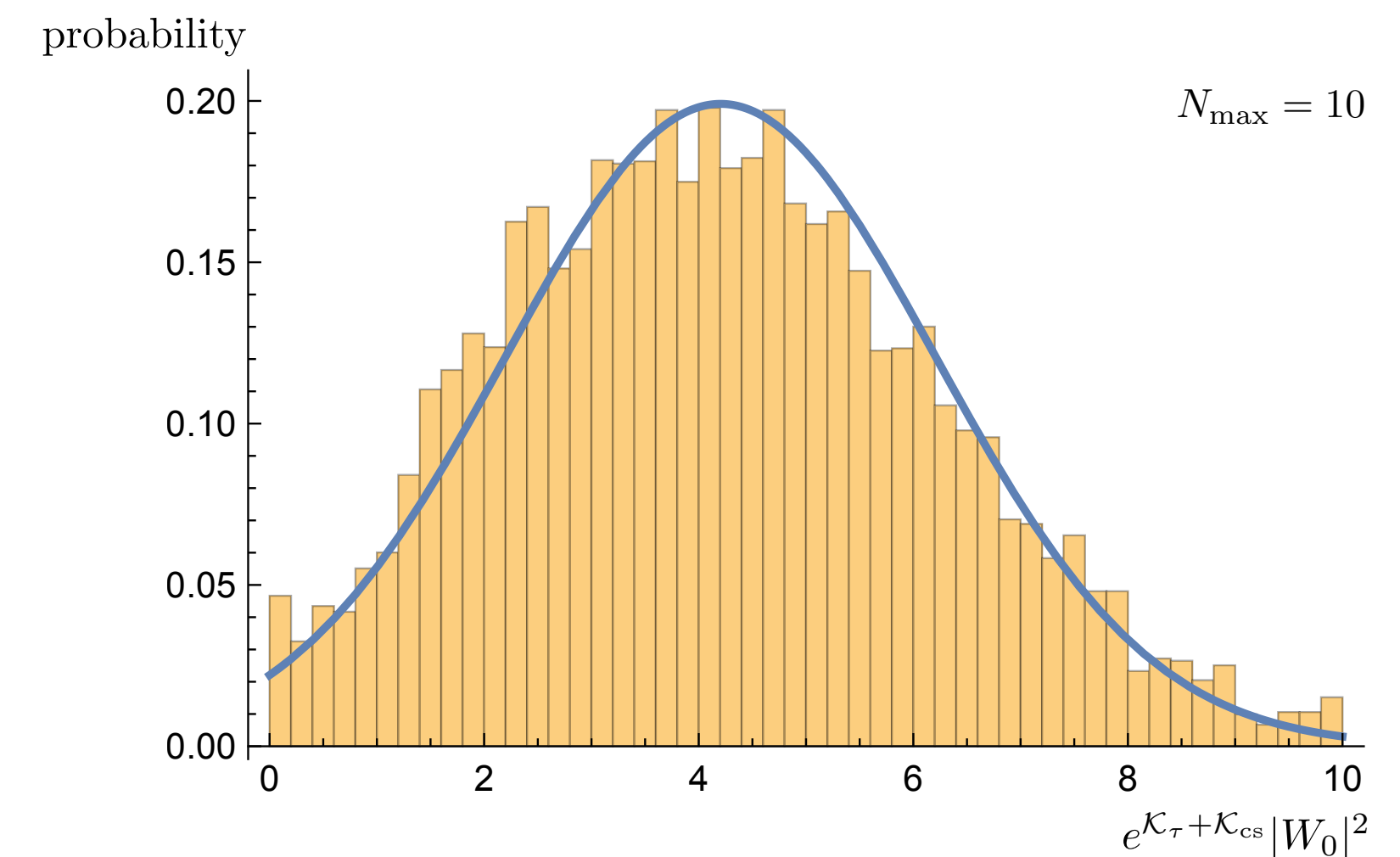


The [gravitino-mass](#) is determined by the expression

$$e^{\mathcal{K}_{\tau} + \mathcal{K}_{cs}} |W_0|^2 .$$

Excluding $W_0 = 0$, it follows a normal distribution with mean and standard deviation

$$(\mu, \sigma) = (0.427 N_{\max}, 0.204 N_{\max}) .$$



Sampling vs. complete scan.

In a **sampling approach**, one often choses flux quanta from a box of length L_{box} as

$$h^I, h_I, f^I, f_I \in [-L_{\text{box}}, +L_{\text{box}}] .$$

The **ratio** $r(L_{\text{box}}, N_{\text{max}})$ of vacua possibly-captured in this way can be defined as

$$\frac{\# \text{ of vacua for box of size } L_{\text{box}}}{\# \text{ of all flux vacua}} \Big|_{N_{\text{flux}} \leq N_{\text{max}}} .$$

L_{box} N_{max}	1	2	3	4	5	10	15	25	35
1	0.667	1	1	1	1	1	1	1	1
2	0.173	0.808	0.962	1	1	1	1	1	1
3	0.051	0.475	0.814	0.898	0.960	1	1	1	1
4	0.019	0.344	0.629	0.838	0.902	0.996	1	1	1
5	0.009	0.211	0.483	0.691	0.826	0.982	1	1	1
6	0.004	0.137	0.391	0.604	0.744	0.968	0.998	1	1
7	0.002	0.087	0.305	0.503	0.659	0.942	0.991	1	1
8	0.001	0.060	0.239	0.443	0.588	0.917	0.983	1	1
9	0.001	0.040	0.199	0.378	0.526	0.886	0.970	0.999	1
10	0.001	0.026	0.156	0.328	0.474	0.860	0.958	0.997	1

The ratio $r(L_{\text{box}}, N_{\text{max}})$.

1. motivation
2. flux vacua
3. the mirror octic
4. results
5. summary

- Summary ::
- We developed a strategy to determine [all](#) flux vacua for type IIB orientifold compactifications.
 - We applied this strategy to the [mirror octic](#) for $N_{\text{flux}} \leq 10$.
 - We analysed properties of the [dataset](#).

- Remarks ::
- The main [technical difficulty](#) is the required computing time.
 - The dataset (together with a Mathematica notebook) is [freely available](#).