

TACHYON INFLATION WARPING AND FLUXES

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Inflation

Scale factor $a \sim e^{Ht}$ $H \sim \text{const.}$

- Slow roll:

$$\epsilon \equiv \frac{M_p^2}{2} \left(\frac{V'}{V} \right)^2 \ll 1$$

$$\eta \equiv M_p^2 \frac{V''}{V} \ll 1$$

- e-folds:

$$N \equiv \int H(t) dt = \frac{1}{M_p^2} \int \frac{V}{V'} d\phi \approx 60$$

- density perturbations

$$\delta_H = \frac{1}{5\pi\sqrt{3}} \frac{V^{3/2}}{M_p V'} = 1.91 \times 10^{-5}$$

- spectral index:

$$n-1 \approx 2\eta - 6\epsilon$$

Inflation from String Theory

Progress:

* Candidate inflatons

- Closed String Moduli BG
- Open String Moduli DT
(D-brane separation)

* Explicit computable potentials

- Racetrack inflation
B-P, BCEGR KLQ
(inflaton: axion partner of volume modulus)

- Brane - Antibrane

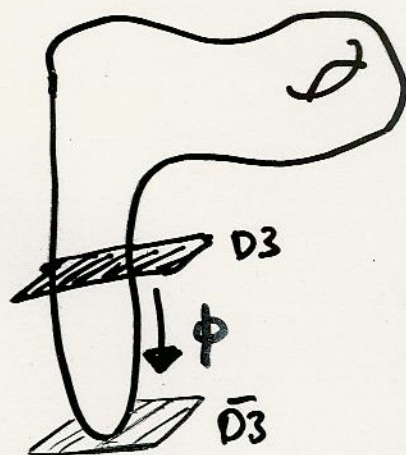
BMNQ²
DSS ... KKLT²
...

(Also D3-D7) KP...



within KKLT scenario

e.g. Brane - Antibrane.



Potential $V(\rho, \phi)$

* η - problem. (... KKLM MT)

$$V_F = e^K \{ \dots \}$$

$$\Rightarrow \eta = \mathcal{J}(1) + \dots \ll 1$$

$\Rightarrow$ fine tuning

($\sim 1/100 - 1/1000$) BCS ϕ , BHK, ...

* Tachyon

• End inflation (hybrid) !
BMNR

• Topological defects

(cosmic strings)
JST, CMP, ...

• Universal for Non-BPS.
Sev.

Tachyon as Inflaton?

A, B, ..., M, ..., K, L

Usually not:

$$V(T) = \frac{V_0}{\cosh(T)}$$

no small parameters
 $\Rightarrow \epsilon, \eta \sim 1$.

But:

$$\mathcal{L}_T = A V(T) \sqrt{1 + B \partial_\mu T \partial^\mu T} \dots R \dots$$

with A, B "warp" constants

$$\epsilon = \frac{M_P^2}{2AB} \left(\frac{V''}{V^3} \right)$$

$$\eta = \frac{M_P^2}{AB} \left[\frac{V'''}{V^2} - \frac{1}{2} \left(\frac{V''}{V^2} \right)^2 \right]$$

$$N = \frac{AB}{M_P^2} \int \frac{V^2}{V'} dT$$

$\Rightarrow AB \gg 1 \Rightarrow$ slow roll inflation?

Also:

$$\delta_H^2 = \frac{A^2 B}{75\pi^2 M_P^2} \left(\frac{V^2}{V_i} \right)^2 \sim 10^{-10}$$

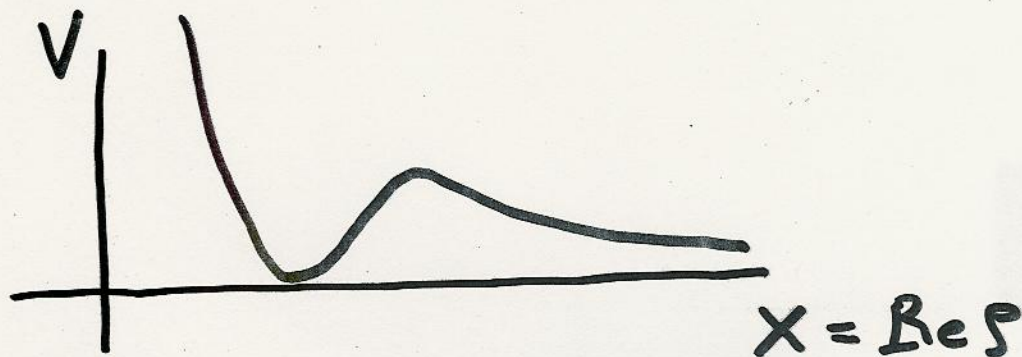
$\Rightarrow$ Choose $\gamma = \frac{m_s}{M_P}$, A, B

to get $N \sim 60$; $\delta_H \sim 10^{-5}$

But γ, A, B are calculable
in concrete scenarios.

Q: Can A, B, γ for Non BPS
D-branes in KKLT Scenario
give rise to inflation?

KKLT:



KKLT + Non-BPS

$$V(x) \rightarrow V(x) + V_{\text{Tachyon}}(x, T)$$

$$\mathcal{L}_{\text{Tachyon}} = A(x, \omega, \underline{F}) \sqrt{1 - B(\omega) \dot{T}^2}$$

$\downarrow$
 x

$$\cdot ds^2 = e^{2\omega - 6\psi + 6\psi_0} ds_{\text{spacetime}}^2 + e^{-2\omega + 2\psi} d\Sigma^2$$

$(e^{4\psi} \equiv X)$

• $F \rightarrow \underline{E}, \underline{B}$ fluxes on Dp

$$A \sim g^{-1} e^{(7-p)\omega_0} h_{p-3} X_0^3 X^{\frac{p-15}{4}} \sqrt{1+B^2} \sqrt{1-E^2}$$

$$B \sim e^{-2\omega} \left(\frac{X}{X_0}\right)^{3/2} (1-E^2)^{-1/2}$$

$$\gamma^2 \equiv \left(\frac{m_J}{M_P}\right)^2 \simeq \frac{g^2 (2\pi)^7}{2 X_0^{3/2}} \frac{e^{2\omega}}{C} \left(\frac{X}{X_0}\right)^{-3/2} (1-E^2)^{1/2}$$

$$h_{p-3} = \int d^6 \xi \sqrt{\tilde{g}}$$

$$C = \int d^6 \xi e^{-4\omega} \sqrt{\tilde{g}}$$

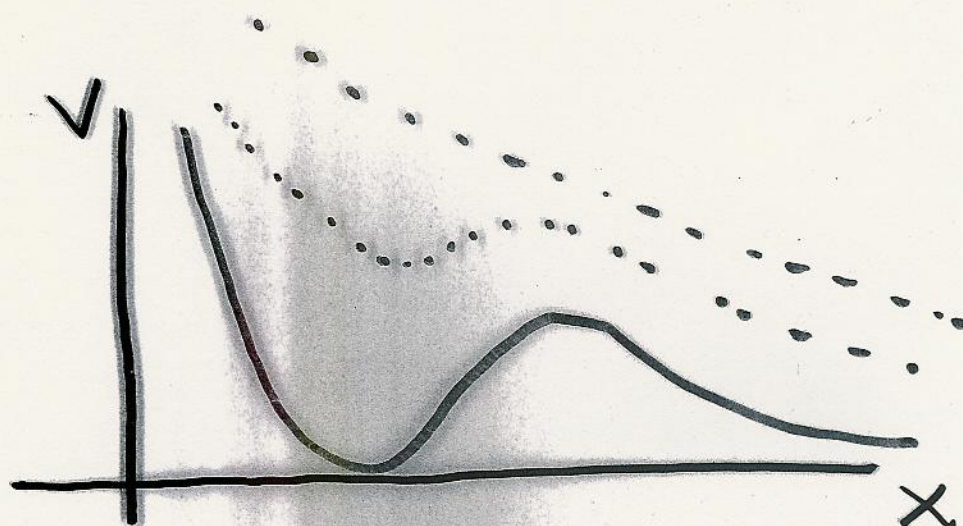
Warp factor $e^{-2\omega}$:

$$\nabla^2 e^{-4\omega} \simeq g_s G_{mnp}^2$$

$\downarrow$
 3-fluxes
 (K, M)

Effects of $V_{tachyon}$:

*



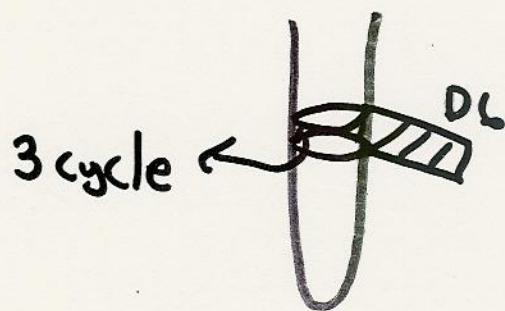
* Slow roll ?

$$(p, e^{\omega}, \underline{B}, \underline{E})$$

$\searrow$
 $K, M + \text{geometry}$

Examples:

(1) $p=6$ on Conifold throat
 $E=B=0$



Tip :

$$ds_7^2 \approx e^{2\omega} ds_4^2 + e^{2\omega} R^2 ds_3^2$$

$$R \approx g_s M$$

$$e^{2\omega} \sim e^{-\frac{2\pi K}{g_s M}}$$

(Thanks: Chen, Razumakovs, Trivedi)

$$N \sim \frac{X_0}{C}^{-3/4} e^{-\frac{2\pi K}{3M}} (3M)^{3/4}$$

$$\text{If } K g_s M \gg 1 \Rightarrow N \sim 10^2$$

but Volume : $C \geq \left(\frac{2\pi}{3}\right)^3 R^6$ (dW, k, v)
constraint

(~ Giant inflaton case)

Improvements:

- $E \neq 0$ $N \rightarrow N/\sqrt{1-E^2}$
 $\Rightarrow E \sim 1 \Rightarrow N \gg 1$
 (but fine tuning)

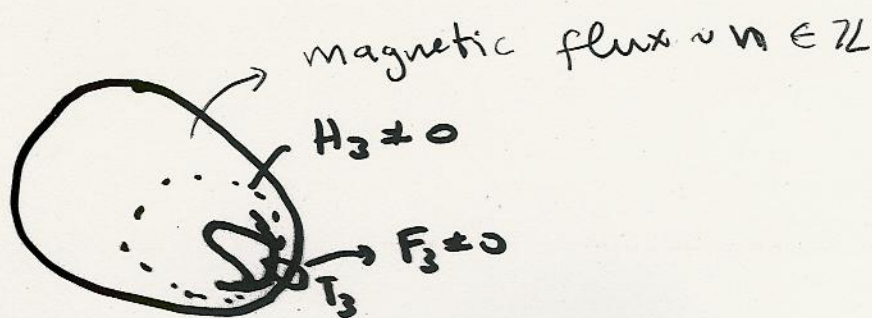
- $B \neq 0$ $P=6$; D6 wrap throat
 (Need 2-cycles)
 $B \gg 1$.

- $P=8$; $B \neq 0$; "middle of throat"

(2.) T_3 fibrations

$$p=6$$

(SYZ)



$$\nabla^2 e^{-4w} \sim g_s \frac{F}{V} \cdot \frac{H}{V_x}$$

$$e^{-2w} \sim g_s^{1/2} \sqrt{\frac{FH}{V_x V}} \quad (x)$$

Constraints: ISD: $F/V_x = \frac{H}{V_x}$

Tadpoles: $FH \lesssim 10^4$?

$$\left. \begin{array}{l} N \sim 10^2 \\ g_s^2 \sim 10^{-10} \end{array} \right\}$$

$$\begin{array}{l} FH \sim 10^4 \\ X_0 \sim 10^4 \\ n \sim 10^8 \end{array}$$



($p=8, \dots$).

Time evolution + Sen's Conjecture

Open string completeness
conjecture:

Non BPS D-brane $\rightarrow$ massive
closed
string
modes.

$\Rightarrow$ back reaction $\Rightarrow$ closed string
description
complicated.

Sen: OSFT describes fully decay
of D-branes (w/o closed strings)
+ Classical results correctly
describe evolution of quantum
ev's (Ehrenfest).

Equations:

$$3H \left(\frac{AB\dot{V}\dot{T}}{\sqrt{1-B\dot{T}^2}} \right) + \left(\dot{\quad} \right) = -A\sqrt{1-B\dot{T}^2} V'$$

$$6 \left(\frac{\dot{X}}{X} \right) + 18H \frac{\dot{X}}{X} + 2X\hat{V}_x + 2X(AV\sqrt{1-B\dot{T}^2})_x = 0$$

$$\frac{\ddot{a}}{a} = -\frac{1}{2} \left(\frac{\dot{X}}{X} \right)^2 + \frac{\hat{V}}{6} + \frac{AV}{6} \frac{1-\frac{3}{2}B\dot{T}^2}{\sqrt{1-B\dot{T}^2}}$$

Sen's
Conjecture:

$$T(t) = T(-t)$$

$$X(t) = X(-t)$$

$$a(t) = a(-t)$$

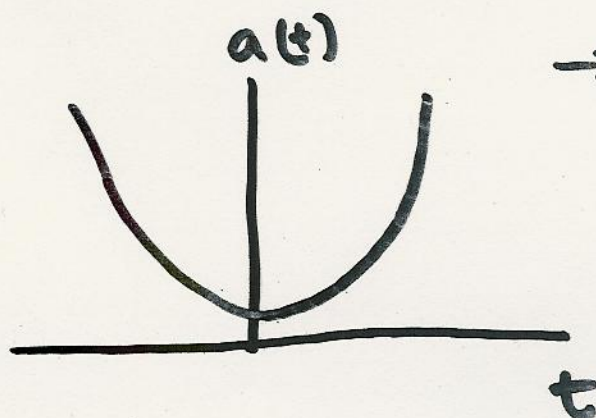
• Initial conditions:

$$\dot{a}(0) = \dot{T}(0) = \dot{X}(0) = 0$$

• Friedmann's: $H^2 = \frac{1}{4} \left(\frac{\dot{X}}{X} \right)^2 + \underbrace{\frac{1}{6}\hat{V} + \frac{1}{6}\frac{AV}{\sqrt{1-B\dot{T}^2}} - \frac{\kappa}{a^2}}_{=0 \text{ at } t=0!}$

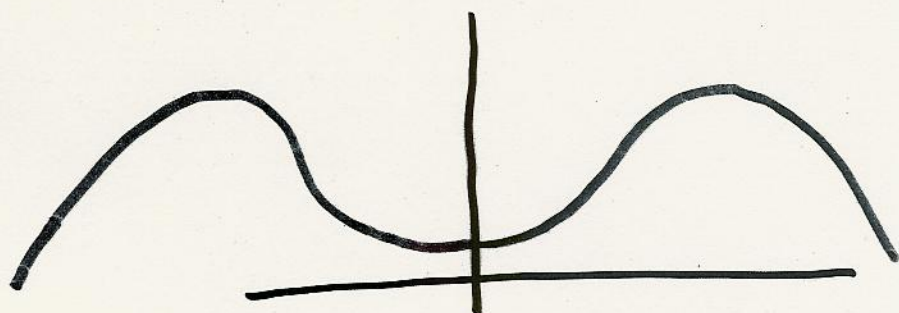
$\Rightarrow$ Free $T(0), X(0).$ ($\kappa=1$).

Time evolution:



→ inflation
+
compactification
+
bounce 4d.

or



big crunch
+
decompact.*
+
big-bang

$$\frac{\ddot{a}}{a} \sim - \left(\frac{\dot{X}}{4X} \right)^2$$

* Need fine-tuning to get inflation + decompactification.

Conclusions

- * General scenario for inflation + moduli fixing
- * 3 - Fluxes
+
magnetic (Wilson Line inflation (in progress))
- * Initial conditions.
(inflation + bounce $\Leftrightarrow$ big crunch decompact.)