

On the Open String Landscape

Fernando Marchesano

University of Wisconsin-Madison

String Phenomenology 2005

LMU München, June 13-18

J. Gomis, F. M., D. Mateos, [hep-th/0506179](https://arxiv.org/abs/hep-th/0506179)

Motivation

Moduli Problem

- Standard Calabi-Yau $\mathcal{N} = 1$ string compactifications suffer from a well-known **problem** in string phenomenology: plenty of **massless scalars**.
- Insight: **Flux** compactifications have **less moduli**

[Dasgupta, Rajesh, Sethi]

[Giddings, Kachru, Polchinski]

Motivation

Moduli Problem

- Standard Calabi-Yau $\mathcal{N} = 1$ string compactifications suffer from a well-known **problem** in string phenomenology: plenty of **massless scalars**.
- Insight: **Flux** compactifications have **less moduli** [DRS,GKP]
- ★ We have a **good understanding** of moduli in **CY** compactifications
- ★ What about $\mathcal{N} = 1$ **$SU(3)$ -structure** compactifications?

... the hope is that they have less moduli (or none).

Moduli and Fluxes

Simplest Example

- Type IIB on a conformal Calabi-Yau, with background 3-form fluxes

$$G_3 = F_3 - \tau H_3 \quad \left\{ \begin{array}{l} F_3 \quad \text{RR flux} \\ H_3 \quad \text{NSNS flux} \\ \tau \quad \text{complex dilaton} \end{array} \right.$$

→ New ingredient: GVW superpotential $W = \int G_3 \wedge \Omega$

Moduli and Fluxes

Simplest Example

- Type IIB on a conformal Calabi-Yau, with background 3-form fluxes

$$G_3 = F_3 - \tau H_3 \quad \begin{cases} F_3 & \text{RR flux} \\ H_3 & \text{NSNS flux} \\ \tau & \text{complex dilaton} \end{cases}$$

→ New ingredient: GVW superpotential $W = \int G_3 \wedge \Omega$

- The existence of W imposes new constraints on complex structure moduli: $DW = W = 0$ [Kachru, Schulz, Trivedi]

→ Reduced moduli space: $\mathcal{M}_{cpx}^{G_3} \subset \mathcal{M}_{cpx}^{CY}$

Moduli and Fluxes

Simplest Example

- Type IIB on a conformal Calabi-Yau, with background 3-form fluxes

$$G_3 = F_3 - \tau H_3 \quad \begin{cases} F_3 & \text{RR flux} \\ H_3 & \text{NSNS flux} \\ \tau & \text{complex dilaton} \end{cases}$$

→ New ingredient: GVW superpotential $W = \int G_3 \wedge \Omega$

- The existence of W imposes new constraints on complex structure moduli: $DW = W = 0$ [Kachru, Schulz, Trivedi]

→ Reduced moduli space: $\mathcal{M}_{cpx}^{G_3} \subset \mathcal{M}_{cpx}^{CY}$

- Equivalently, $\mathcal{N} = 1$ conditions read G_3 harmonic (2,1)-form

→ Given an $\mathcal{N} = 1$ vacuum, only by moving in $\mathcal{M}_{cpx}^{G_3} \subset \mathcal{M}_{cpx}^{CY}$ this condition will not be spoiled

Fluxes and D-branes

What about D-branes?

- We know by some partial results that the moduli space of D-branes can also be affected by G_3
- Examples:
 - D7-branes
 - [Görllich, Kachru, Tripathy, Trivedi]
 - [Cascales, Uranga]
 - [Cámara, Ibáñez, Uranga]
 - [Lüst, Mayr, Reffert, Stieberger]
 - Euclidean D3-branes
 - [Tripathy, Trivedi]
 - [Saulina]
 - [Kallosh, Kashani-Poor, Tomasiello]

Fluxes and D-branes

What about D-branes?

- We know by some partial results that the moduli space of D-branes can also be affected by G_3

- Examples:

- D7-branes

- [Görllich, Kachru, Tripathy, Trivedi]
[Cascales, Uranga]
[Cámara, Ibáñez, Uranga]
[Lüst, Mayr, Reffert, Stieberger]

- Euclidean D3-branes

- [Tripathy, Trivedi]
[Saulina]
[Kallosh, Kashani-Poor, Tomasiello]

... but we don't have a complete and general understanding of these facts

Fluxes and D-branes

What about D-branes?

- We know by some partial results that the moduli space of D-branes can also be affected by G_3
- Examples:
 - D7-branes [GKTT, CU, CIU, LMRS]
 - Euclidean D3-branes [TT, Saulina, KKT]

IDEA

Implement the same approach
for Open Strings
as done for Closed Strings

D-branes and Moduli

- First approach: F/M-theory lift + GVW superp. $W = \int G_4 \wedge \Omega$

But...

- Needs global knowledge of $SU(4)$ -holonomy geometries
- D-brane intuition is lost

D-branes and Moduli

- First approach: F/M-theory lift + GVW superp. $W = \int G_4 \wedge \Omega$
- **Second approach:** Type IIB background + D-brane probe approx.
 - Write down SUSY conditions for D-branes
 - Find the D-brane deformations that do not spoil them

D-branes and Moduli

- First approach: F/M-theory lift + GVW superp. $W = \int G_4 \wedge \Omega$
- **Second approach:** Type IIB background + D-brane probe approx.
 - Write down **SUSY conditions** for D-branes
 - Find the D-brane **deformations** that do not spoil them

In the **conformal CY case**, we hope to find a description in terms of

$$\left. \begin{array}{l} \text{Old moduli space} \\ \text{New restrictions} \end{array} \right\} \Rightarrow \text{Reduced moduli space}$$

D-branes & SUSY

- Which are the supersymmetric D-branes in flux compactifications?
 - Possible approach: generalized calibrations [Cascales, Uranga]

D-branes & SUSY

- Which are the **supersymmetric D-branes** in **flux compactifications**?
 - Possible approach: **generalized calibrations** [Cascales, Uranga]
 - We will use another technique: **κ -symmetry**

$$\Gamma \epsilon = \epsilon$$

ϵ : cov. constant spinor of the background

Γ : encodes the D-brane embedding

SUSY & κ -symmetry

- Previous analysis of Mariño, Minasian, Moore, and Strominger
 - Calabi-Yau background \mathcal{M}_6
 - Euclidean $D(2n - 1)$ -branes wrapping $\mathcal{S}_{2n} \subset \mathcal{M}_6$
 - $U(1)$ gauge bundle \mathcal{F} in \mathcal{S}_{2n}

SUSY & κ -symmetry

- Previous analysis of Mariño, Minasian, Moore, and Strominger
 - Calabi-Yau background \mathcal{M}_6
 - Euclidean D(2n - 1)-branes wrapping $\mathcal{S}_{2n} \subset \mathcal{M}_6$
 - U(1) gauge bundle \mathcal{F} in \mathcal{S}_{2n}

Result:

$$e^{iJ-\mathcal{F}}|_{\mathcal{S}_{2n}}^{\text{top}} = e^{i\theta} \sqrt{|g^{\mathcal{S}_{2n}} + \mathcal{F}|} \frac{d\text{vol}_{\mathcal{S}_{2n}}}{\sqrt{|g^{\mathcal{S}_{2n}}|}}$$

$$\iota_X \Omega \wedge e^{iJ-\mathcal{F}}|_{\mathcal{S}_{2n}}^{\text{top}} = 0, \quad \forall X \in T\mathcal{M}_6$$

Ω : Holomorphic 3-form J : Kähler 2-form $\mathcal{F} = 2\pi\alpha' F + B_2|_{\mathcal{S}_{2n}}$

θ : $\mathcal{N} = 1 \subset \mathcal{N} = 2$ parameter $(\iota_X C_{p+1})_{a_1 \dots a_p} = X^\rho (C_{p+1})_{\rho a_1 \dots a_p}$

κ -symmetry & Fluxes

- One can extend the **MMMS** computation to **more general backgrounds**, since the main ingredients needed are
 - **Hermitian** metric
 - p -forms Ω , J defined as **spinor bilinears**
 - 2 **SUSY generators** $\eta_+^* = \eta_-$

κ -symmetry & Fluxes

• One can extend the **MMMS** computation to **more general backgrounds**, since the main ingredients needed are

- Hermitian metric
- p -forms Ω , J defined as spinor bilinears
- 2 SUSY generators $\eta_+^* = \eta_-$

⇒ Conditions satisfied for any $SU(3)$ -structure manifold

[Graña, Minasian, Petrini, Tomasiello]



MMMS equations still apply

(Warning! Now $d\Omega$, dJ , $d\mathcal{F} \neq 0$)

Fluxes & D-branes

Simplest Example

- Warped Calabi-Yau with ISD G_3 fluxes

$$ds^2 = \Delta(y)^{-1} \eta_{\mu\nu} dx^\mu dx^\nu + \Delta(y) g_{mn}^{\text{CY}} dy^m dy^n,$$

$$\Omega = \Delta^{3/2} \Omega^{\text{CY}} \quad J = \Delta J^{\text{CY}}$$

Fluxes & D-branes

Simplest Example

- Warped Calabi-Yau with ISD G_3 fluxes

$$ds^2 = \Delta(y)^{-1} \eta_{\mu\nu} dx^\mu dx^\nu + \Delta(y) g_{mn}^{\text{CY}} dy^m dy^n,$$

$$\Omega = \Delta^{3/2} \Omega^{\text{CY}} \quad J = \Delta J^{\text{CY}}$$

MMMS Eqs:

$$e^{iJ - \mathcal{F}}|_{S_{2n}}^{\text{top}} = - \sqrt{|g^{S_{2n}} + \mathcal{F}|} \frac{d\text{vol}_{S_{2n}}}{\sqrt{|g^{S_{2n}}|}}$$
$$\iota_X \Omega \wedge e^{iJ - \mathcal{F}}|_{S_{2n}}^{\text{top}} = 0, \quad \forall X \in TM_6$$

Fluxes & D-branes

Simplest Example

- Warped Calabi-Yau with ISD G_3 fluxes

$$ds^2 = \Delta(y)^{-1} \eta_{\mu\nu} dx^\mu dx^\nu + \Delta(y) g_{mn}^{\text{CY}} dy^m dy^n,$$

$$\Omega = \Delta^{3/2} \Omega^{\text{CY}} \quad J = \Delta J^{\text{CY}}$$

MMMS Eqs:

$$e^{iJ - \mathcal{F}} \Big|_{\mathcal{S}_{2n}}^{\text{top}} = - \sqrt{|g^{\mathcal{S}_{2n}} + \mathcal{F}|} \frac{d\text{vol}_{\mathcal{S}_{2n}}}{\sqrt{|g^{\mathcal{S}_{2n}}|}}$$

\mathcal{S}_{2n} holomorphic $\mathcal{F}^{(2,0)} = 0$

Fluxes & D-branes

Simplest Example

- Warped Calabi-Yau with ISD G_3 fluxes

$$ds^2 = \Delta(y)^{-1} \eta_{\mu\nu} dx^\mu dx^\nu + \Delta(y) g_{mn}^{\text{CY}} dy^m dy^n,$$

$$\Omega = \Delta^{3/2} \Omega^{\text{CY}} \quad J = \Delta J^{\text{CY}}$$

MMMS Eqs:

$$J|_{S_2} + i\mathcal{F} = i\sqrt{|g^{S_2} + \mathcal{F}|} \frac{d\text{vol}_{S_2}}{\sqrt{|g^{S_2}|}}$$

S_2 holomorphic

- $n = 1 \rightarrow$ D5-branes or Euclidean D1's

$$J|_{S_2} \rightarrow 0 \quad \Rightarrow \quad \text{fractional D3-brane}$$

Fluxes & D-branes

Simplest Example

- Warped Calabi-Yau with ISD G_3 fluxes

$$ds^2 = \Delta(y)^{-1} \eta_{\mu\nu} dx^\mu dx^\nu + \Delta(y) g_{mn}^{\text{CY}} dy^m dy^n,$$

$$\Omega = \Delta^{3/2} \Omega^{\text{CY}} \quad J = \Delta J^{\text{CY}}$$

MMMS Eqs:

$$\begin{aligned} J|_{\mathcal{S}_4} \wedge \mathcal{F} &= 0 \\ \mathcal{S}_4 \text{ holomorphic} \quad \mathcal{F}^{(2,0)} &= 0 \end{aligned}$$

- $n = 2 \rightarrow$ D7-branes or Euclidean D3's

$$\mathcal{F} \text{ is } (1,1) \text{ primitive} \quad \Rightarrow \quad *_4 \mathcal{F} = -\mathcal{F}$$

D3-brane induced charge

Fluxes & D-branes

Simplest Example

- Warped Calabi-Yau with ISD G_3 fluxes

$$ds^2 = \Delta(y)^{-1} \eta_{\mu\nu} dx^\mu dx^\nu + \Delta(y) g_{mn}^{\text{CY}} dy^m dy^n,$$

$$\Omega = \Delta^{3/2} \Omega^{\text{CY}} \quad J = \Delta J^{\text{CY}}$$

MMMS Eqs:

$$\begin{aligned} \frac{1}{3!} J^3 &= \frac{1}{2!} J \wedge \mathcal{F} \\ \mathcal{F}^{(2,0)} &= 0 \end{aligned}$$

- $n = 3 \rightarrow$ D9-branes or Euclidean D5's

$d\mathcal{F} = H_3$ is not defined globally \Rightarrow Freed-Witten anomaly

We need extra sources of \mathcal{F}

D7-branes

- Recall that

$$\mathcal{F} = 2\pi\alpha'F + B_2|_{\mathcal{S}_4} \left\{ \begin{array}{l} F = dA \in H^2(\mathcal{S}_4, \mathbb{Z}) \\ dB_2 = H_3 \end{array} \right\} \Rightarrow d\mathcal{F} = H_3|_{\mathcal{S}_4}$$

→ \mathcal{F} is well-defined only if $[H_3|_{\mathcal{S}_4}] \in H_3(\mathcal{S}_4, \mathbb{Z})$ vanishes

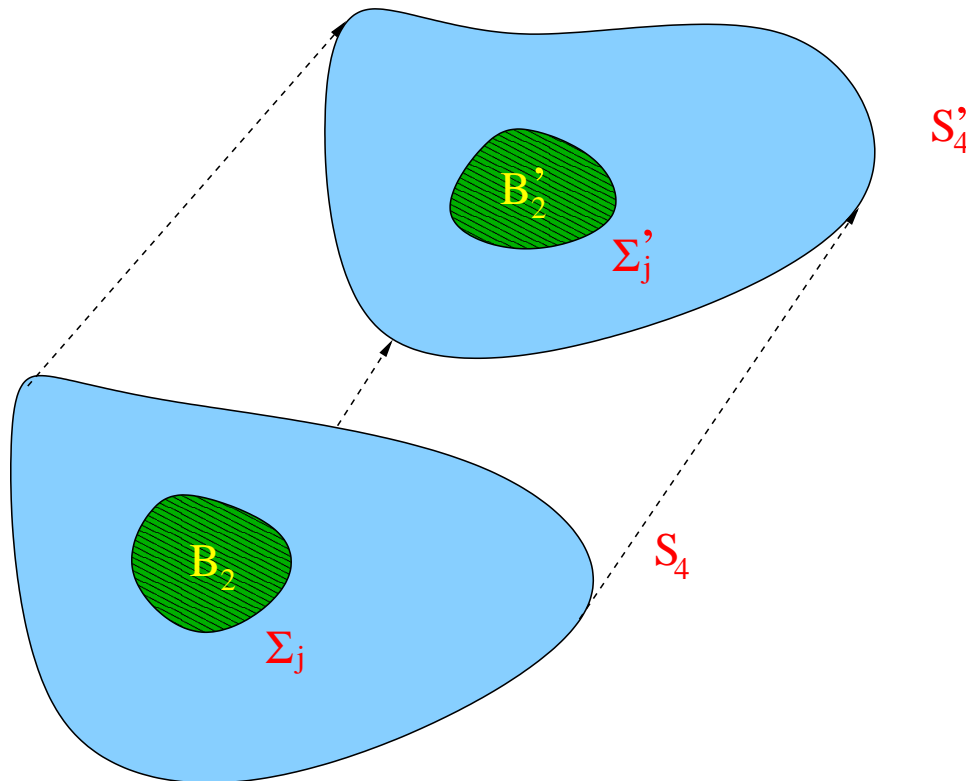
D7-branes

- Recall that

$$\mathcal{F} = 2\pi\alpha'F + B_2|_{\mathcal{S}_4} \left\{ \begin{array}{l} F = dA \in H^2(\mathcal{S}_4, \mathbb{Z}) \\ dB_2 = H_3 \end{array} \right\} \Rightarrow d\mathcal{F} = H_3|_{\mathcal{S}_4}$$

→ \mathcal{F} is **well-defined** only if $[H_3|_{\mathcal{S}_4}] \in H_3(\mathcal{S}_4, \mathbb{Z})$ **vanishes**

- Since B_2 is non-closed, its **pull-back depends on** the 4-cycle \mathcal{S}_4 that the D7-brane is wrapping



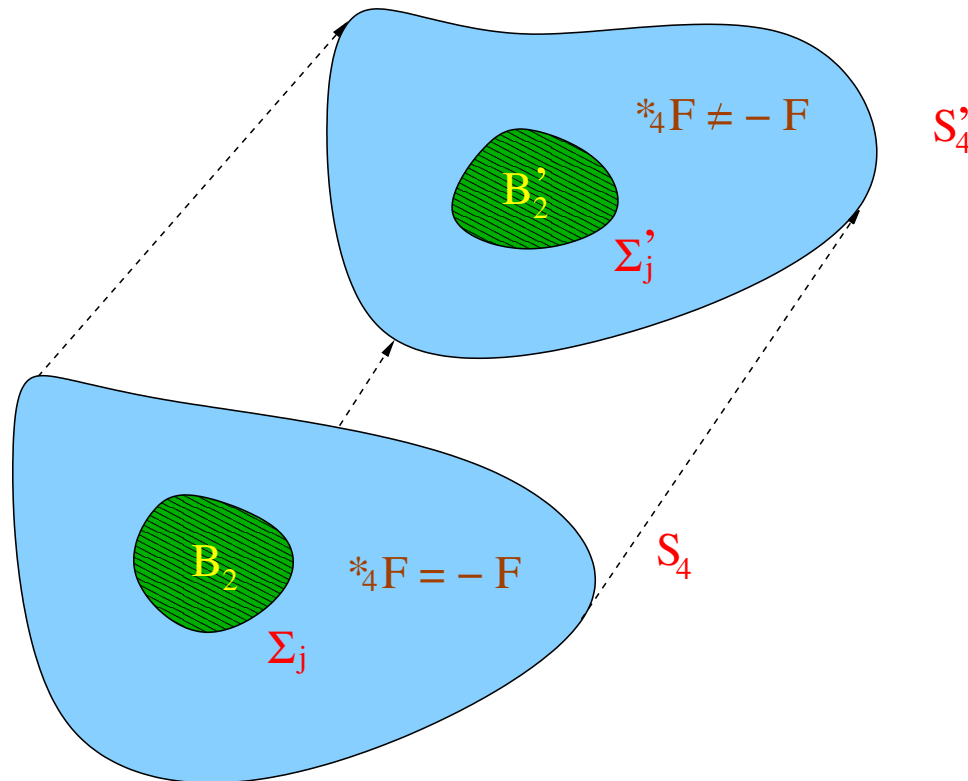
D7-branes

- Recall that

$$\mathcal{F} = 2\pi\alpha'F + B_2|_{\mathcal{S}_4} \left\{ \begin{array}{l} F = dA \in H^2(\mathcal{S}_4, \mathbb{Z}) \\ dB_2 = H_3 \end{array} \right\} \Rightarrow d\mathcal{F} = H_3|_{\mathcal{S}_4}$$

→ \mathcal{F} is well-defined only if $[H_3|_{\mathcal{S}_4}] \in H_3(\mathcal{S}_4, \mathbb{Z})$ vanishes

- A deformation $\mathcal{S}_4 \rightarrow \mathcal{S}'_4$ may not preserve the supersymmetry condition $*_4\mathcal{F} = -\mathcal{F}$



D7-branes II

- In absence of H_3 flux, the moduli space of a D7-brane is given by

Geometric moduli : ζ^a $a = 1, \dots, H^{(0,2)}(\mathcal{S}_4)$

Wilson line moduli : ξ^b $b = 1, \dots, H^{(0,1)}(\mathcal{S}_4)$

[Jockers, Louis]

D7-branes II

- In absence of H_3 flux, the moduli space of a D7-brane is given by

Geometric moduli : ζ^a $a = 1, \dots, H^{(0,2)}(\mathcal{S}_4)$

Wilson line moduli : ξ^b $b = 1, \dots, H^{(0,1)}(\mathcal{S}_4)$

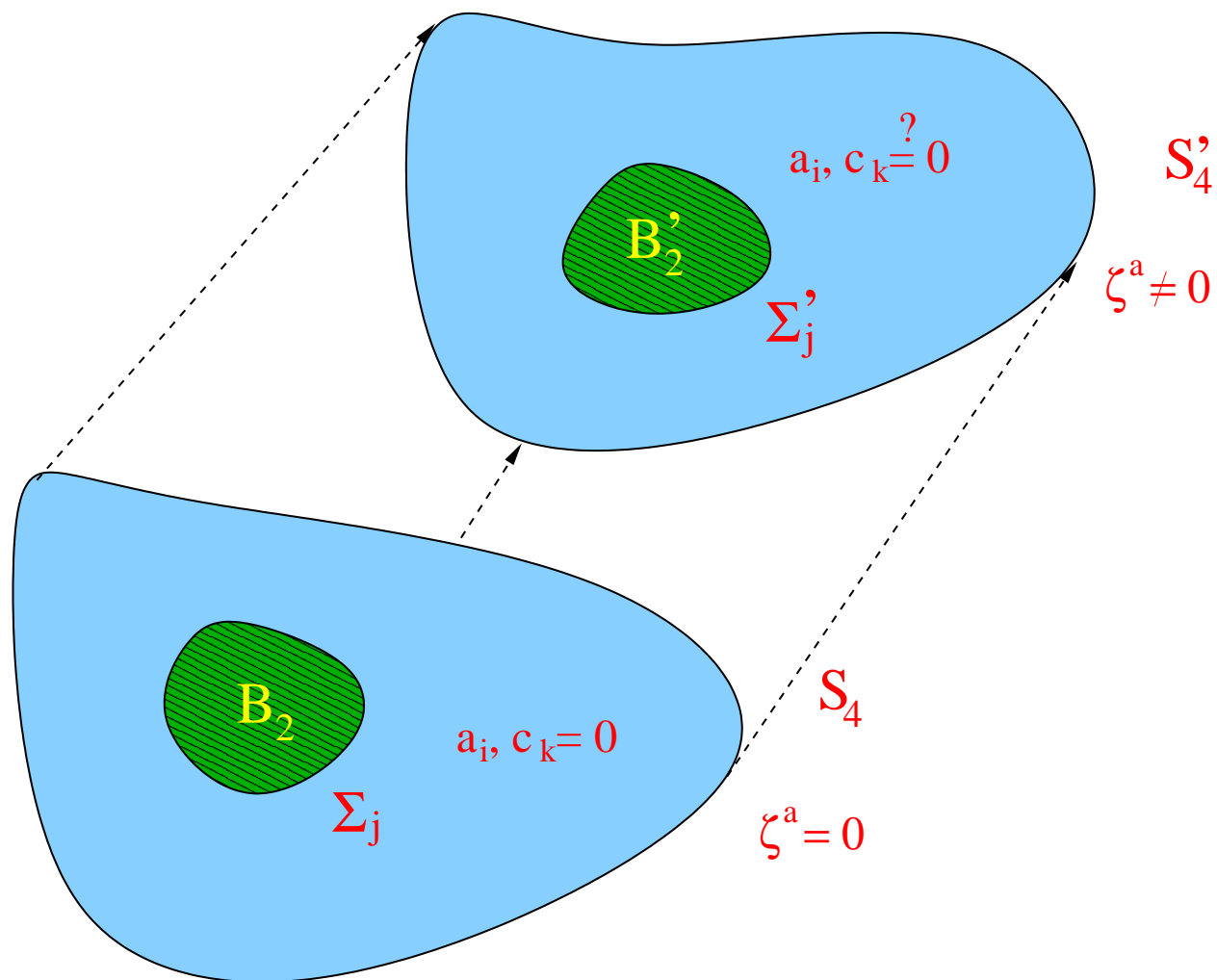
[Jockers, Louis]

- The harmonic part of \mathcal{F} has the form

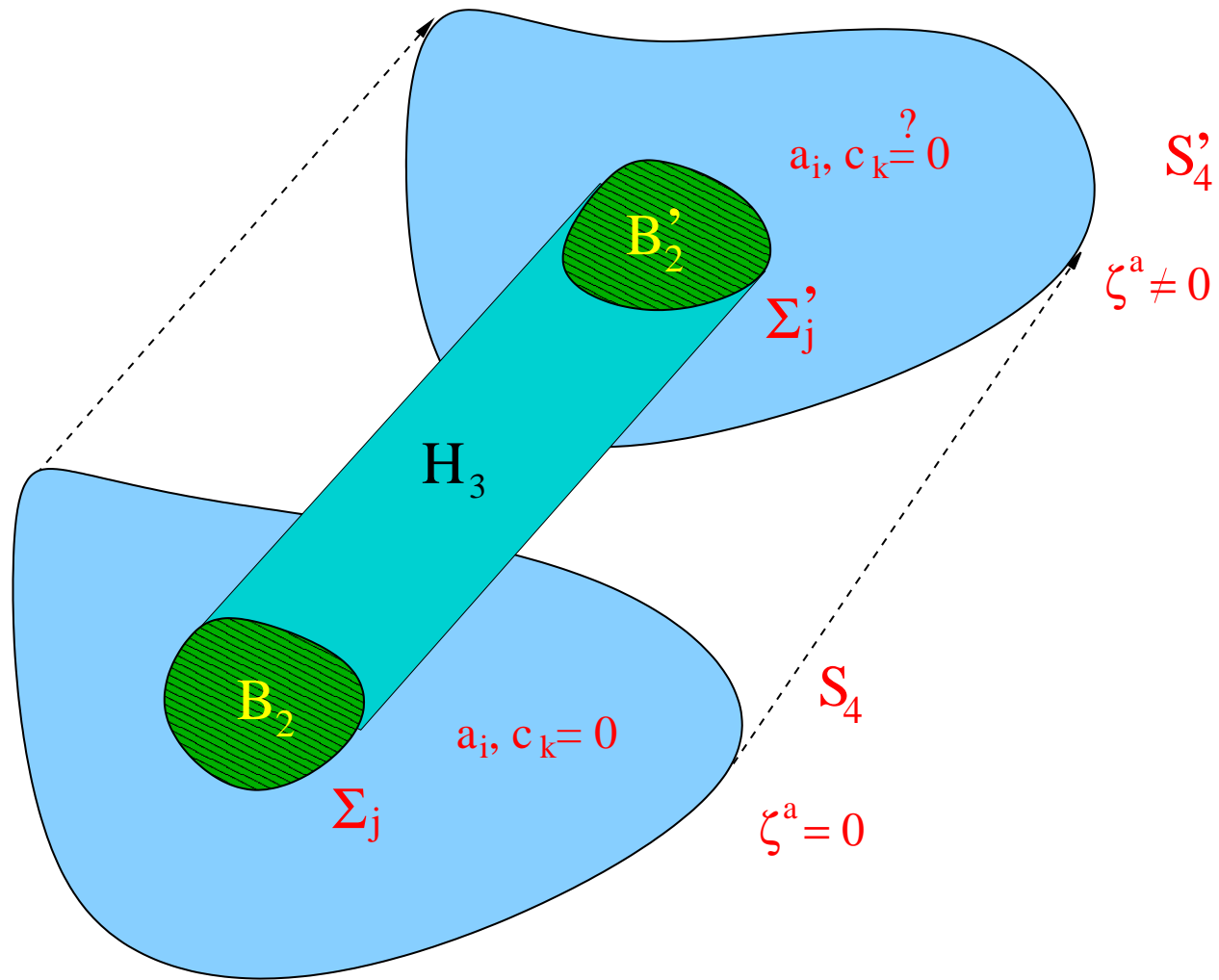
$$\begin{aligned} \mathcal{F}^{har} = & \sum_{i=1}^{h^{2,0}} a_i(\{\zeta^a\}) \alpha^i + \sum_{j=1}^{h_p^{1,1}} b_j(\{\zeta^a\}) \beta^j \\ & + \sum_{k=1}^{h_{np}^{1,1}} c_k(\{\zeta^a\}) \gamma^k + \text{c.c.} \end{aligned}$$

$$\begin{aligned} \bar{\alpha}^i \in \mathcal{H}^{2,0} & \quad \beta^j \in \mathcal{H}_p^{1,1} & \alpha^i \in \mathcal{H}^{0,2} \\ \gamma^k \in \mathcal{H}_{np}^{1,1} & & \end{aligned}$$

D7-branes II



D7-branes II



D7-branes II

$$\mathcal{F}^{har} = \sum_{i=1}^{h^{2,0}} a_i(\{\zeta^a\}) \alpha^i + \sum_{j=1}^{h_p^{1,1}} b_j(\{\zeta^a\}) \beta^j \\ + \sum_{k=1}^{h_{np}^{1,1}} c_k(\{\zeta^a\}) \gamma^k + \text{c.c.}$$

$$\bar{\alpha}^i \in \mathcal{H}^{2,0} \quad \beta^j \in \mathcal{H}_p^{1,1} \quad \alpha^i \in \mathcal{H}^{0,2} \\ \gamma^k \in \mathcal{H}_{np}^{1,1}$$

D7-branes II

$$\mathcal{F}^{har} = \sum_{i=1}^{h^{2,0}} a_i(\{\zeta^a\}) \alpha^i + \sum_{j=1}^{h_p^{1,1}} b_j(\{\zeta^a\}) \beta^j \\ + c_J(\{\zeta^a\}) J^{\text{CY}}|_{\mathcal{S}_4} + \text{c.c.}$$

$$\bar{\alpha}^i \in \mathcal{H}^{2,0} \quad \beta^j \in \mathcal{H}_p^{1,1} \quad \alpha^i \in \mathcal{H}^{0,2} \\ [J|_{\mathcal{S}_4}] \in H_{np}^{1,1}$$

D7-branes II

$$\mathcal{F}^{har} = \sum_{i=1}^{h^{2,0}} a_i(\{\zeta^a\}) \alpha^i + \sum_{j=1}^{h_p^{1,1}} b_j(\{\zeta^a\}) \beta^j \\ + c_J(\{\zeta^a\}) J^{CY}|_{\mathcal{S}_4} + \text{c.c.}$$

$$\bar{\alpha}^i \in \mathcal{H}^{2,0} \quad \beta^j \in \mathcal{H}_p^{1,1} \quad \alpha^i \in \mathcal{H}^{0,2} \\ [J|_{\mathcal{S}_4}] \in H_{np}^{1,1}$$

- Supersymmetry of the background implies

$$\rightarrow G_3 \wedge J = 0 \quad \Rightarrow \quad c_J = \text{const.}$$

$$\rightarrow G_3 \text{ is } (2,1) \quad \Rightarrow \quad a_i = \text{holomorphic on } \{\zeta^a\}$$

D7-branes III

D7-brane moduli

- Old moduli space:

$$\mathcal{M}(D7) = \begin{cases} \text{Geom. deform.} & \rightarrow \mathcal{M}(\mathcal{S}_4) = \{\zeta^i\} & \dim = h^{0,2}(\mathcal{S}_4) \\ \text{Wilson lines} & \rightarrow \mathcal{M}(\mathcal{F}) = \{\xi^j\} & \dim = h^{0,1}(\mathcal{S}_4) \end{cases}$$

D7-branes III

D7-brane moduli

- Old moduli space:

$$\mathcal{M}(D7) = \begin{cases} \text{Geom. deform.} & \rightarrow \mathcal{M}(\mathcal{S}_4) = \{\zeta^i\} & \dim = h^{0,2}(\mathcal{S}_4) \\ \text{Wilson lines} & \rightarrow \mathcal{M}(\mathcal{F}) = \{\xi^j\} & \dim = h^{0,1}(\mathcal{S}_4) \end{cases}$$

- \mathcal{F} is now coupled to $\mathcal{M}(\mathcal{S}_4)$ via a non-trivial H_3 ,
so **supersymmetry restricts** the possible deformations $\{\zeta^i\}$

$$\begin{aligned} a_1(\zeta^1, \dots, \zeta^{h^{0,2}}) &= 0, \\ &\vdots \\ a_{h^{0,2}}(\zeta^1, \dots, \zeta^{h^{0,2}}) &= 0, \end{aligned}$$

D7-branes III

D7-brane moduli

- Old moduli space:

$$\mathcal{M}(D7) = \begin{cases} \text{Geom. deform.} & \rightarrow \mathcal{M}(\mathcal{S}_4) = \{\zeta^i\} & \dim = h^{0,2}(\mathcal{S}_4) \\ \text{Wilson lines} & \rightarrow \mathcal{M}(\mathcal{F}) = \{\xi^j\} & \dim = h^{0,1}(\mathcal{S}_4) \end{cases}$$

- \mathcal{F} is now coupled to $\mathcal{M}(\mathcal{S}_4)$ via a non-trivial H_3 ,
so **supersymmetry restricts** the possible deformations $\{\zeta^i\}$

$$\begin{aligned} a_1(\zeta^1, \dots, \zeta^{h^{0,2}}) &= 0, \\ &\vdots \\ a_{h^{0,2}}(\zeta^1, \dots, \zeta^{h^{0,2}}) &= 0, \end{aligned}$$

→ Reduced moduli space: subvariety $\mathcal{M}^{G_3}(\mathcal{S}_4) \subset \mathcal{M}(\mathcal{S}_4)$,
typically **zero dimensional**

→ Wilson lines untouched

So far so good. . .

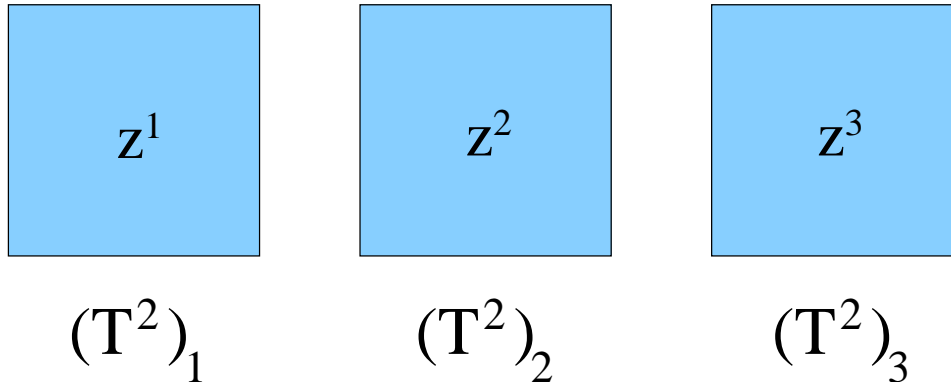
- We are interested in knowing which is the **moduli space of D-branes** in flux compactifications
- In order to work it out, we have deduced the **supersymmetry conditions** for D-branes in manifolds with $SU(3)$ structure, via κ -symmetry

So far so good. . .

- We are interested in knowing which is the **moduli space of D-branes** in flux compactifications
- In order to work it out, we have deduced the **supersymmetry conditions** for D-branes in manifolds with $SU(3)$ structure, via κ -symmetry
- For a **D7-brane wrapping** a divisor \mathcal{S}_4 , we have discovered that \mathcal{F} **depends** non-trivially **on** the space of deformations $\mathcal{M}(\mathcal{S}_4)$
- Most of these **deformations spoil** the **supersymmetry** conditions, which effectively reduces the moduli space to $\mathcal{M}^{G_3}(\mathcal{S}_4) \subset \mathcal{M}(\mathcal{S}_4)$
- $\mathcal{M}^{G_3}(\mathcal{S}_4)$ is the solution to $h^{0,2}$ equations with $h^{0,2}$ unknowns
→ **All geometrical moduli** are typically **lifted**

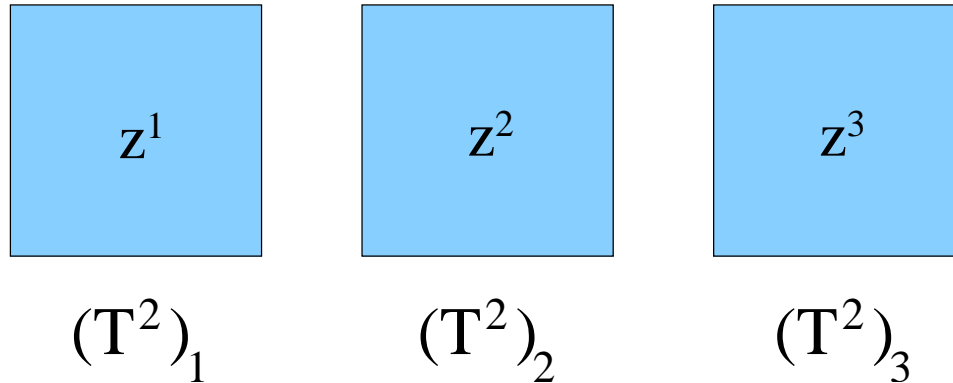
A simple example

- Let us consider a toroidal compactification



A simple example

- Let us consider a toroidal compactification



with constant fluxes

$$F_3 = 4\pi^2\alpha' N (dx^1 \wedge dx^2 \wedge dy^3 + dy^1 \wedge dy^2 \wedge dy^3)$$
$$H_3 = 4\pi^2\alpha' N (dx^1 \wedge dx^2 \wedge dx^3 + dy^1 \wedge dy^2 \wedge dx^3)$$

which create a superpotential

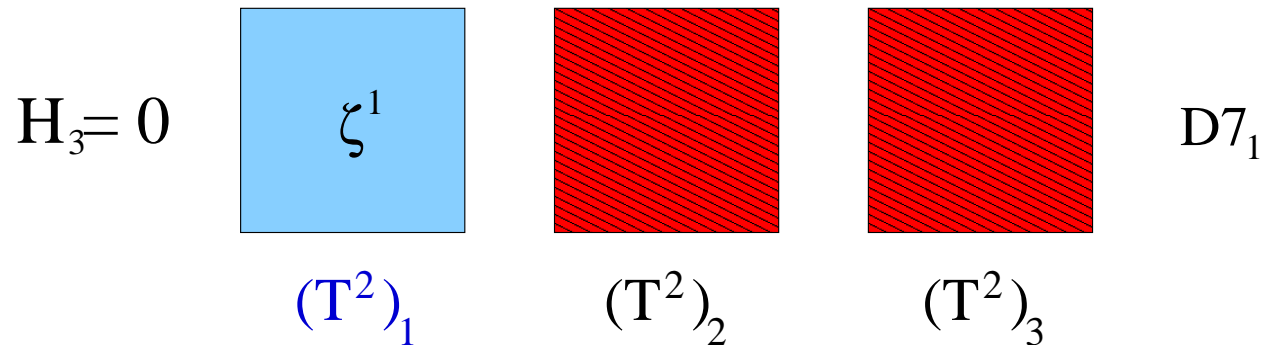
$$W = \int \Omega \wedge G_3 \propto (1 + \tau_1\tau_2) \cdot (1 + \tau_3\tau),$$

\Downarrow

$$\tau_1\tau_2 = -1, \quad \tau_3\tau = -1$$

A simple example

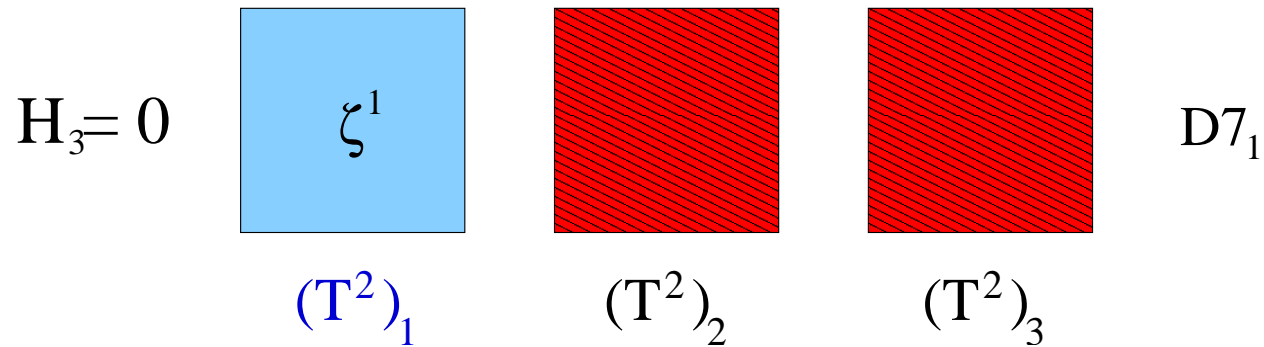
- Let us consider a toroidal compactification



with a $D7_1$ -brane, $h^{0,2}(D7_1) = 1$, $\mathcal{M}(D7_1) = (T^2)_1$

A simple example

- Let us consider a toroidal compactification



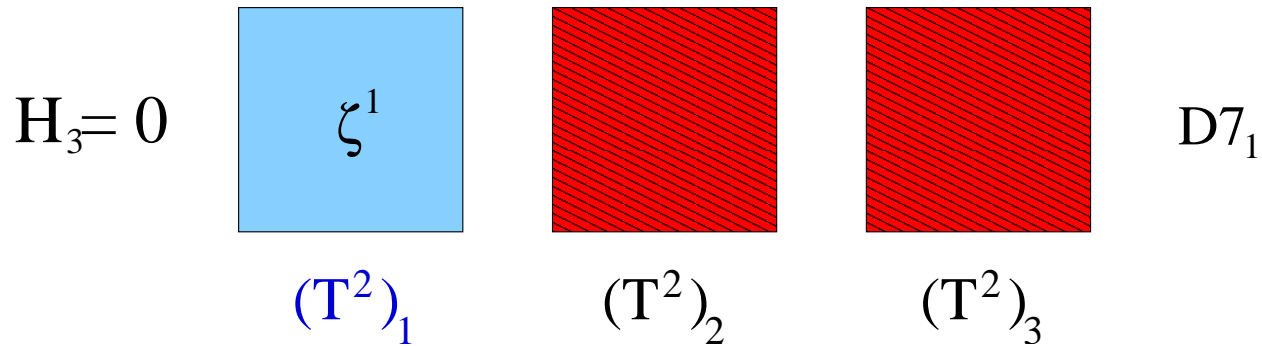
with a $D7_1$ -brane, $h^{0,2}(D7_1) = 1$, $\mathcal{M}(D7_1) = (T^2)_1$

$$H_3 = 4\pi^2\alpha' N (dx^1 \wedge dx^2 \wedge dx^3 + dy^1 \wedge dy^2 \wedge dx^3)$$

$$B_2 = 4\pi^2\alpha' N (x^1 dx^2 \wedge dx^3 + y^1 dy^2 \wedge dx^3)$$

A simple example

- Let us consider a toroidal compactification



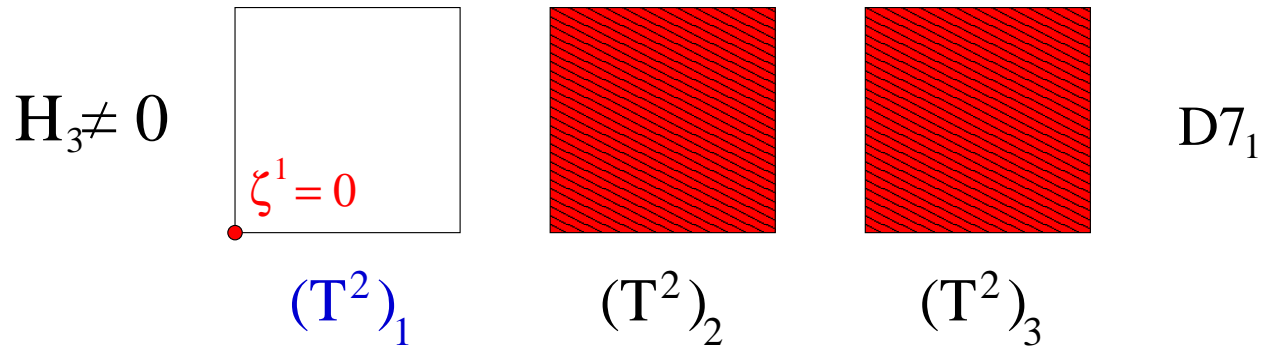
with a $D7_1$ -brane, $h^{0,2}(D7_1) = 1$, $\mathcal{M}(D7_1) = (T^2)_1$

$$H_3 = 4\pi^2\alpha' N (dx^1 \wedge dx^2 \wedge dx^3 + dy^1 \wedge dy^2 \wedge dx^3)$$

$$B_2|_{D7_1} = 4\pi^2\alpha' N \left(\frac{\tau_2}{\tau_2 - \bar{\tau}_2} \zeta^1 d\bar{z}^2 + \frac{\tau_1}{\tau_1 - \bar{\tau}_1} \bar{\zeta}^1 dz^2 \right) \wedge \frac{\text{Im}(\tau_3 d\bar{z}^3)}{\text{Im} \tau^3}$$

A simple example

- Let us consider a toroidal compactification



with a $D7_1$ -brane, $h^{0,2}(D7_1) = 1$, $\mathcal{M}(D7_1) = (T^2)_1$

$$H_3 = 4\pi^2\alpha' N (dx^1 \wedge dx^2 \wedge dx^3 + dy^1 \wedge dy^2 \wedge dx^3)$$

$$B_2|_{D7_1} = 4\pi^2\alpha' N \left(\frac{\tau_2}{\tau_2 - \bar{\tau}_2} \zeta^1 dz^{\bar{2}} + \frac{\tau_1}{\tau_1 - \bar{\tau}_1} \bar{\zeta}^{\bar{1}} dz^2 \right) \wedge \frac{\text{Im}(\tau_3 dz^{\bar{3}})}{\text{Im} \tau^3}$$

↓

$$a_1(\zeta^1) = N\zeta^1\tau_2\tau_3 \quad \Rightarrow \quad \zeta^1 = 0 \quad (\Rightarrow B_2|_{D7_1} = 0)$$

Finding the discretum

- Are we actually finding all the supersymmetric solutions?

Finding the discretum

- Are we actually finding all the supersymmetric solutions?
- In general we have to consider

$$\mathcal{F}_{D7_1} = 2\pi\alpha' F_{D7_1} + B_2|_{D7_1}$$

$$\begin{aligned} B|_{D7_1} &= 4\pi^2\alpha' N \left(\zeta_x^1 dx^2 \wedge dx^3 + \zeta_y^1 dy^2 \wedge dx^3 \right) \\ F_{D7_1} &= 2\pi \left(n_1 dx^2 \wedge dx^3 + n_2 dy^2 \wedge dx^3 + \dots \right) \end{aligned}$$

Finding the discretum

- Are we actually finding all the supersymmetric solutions?
- In general we have to consider

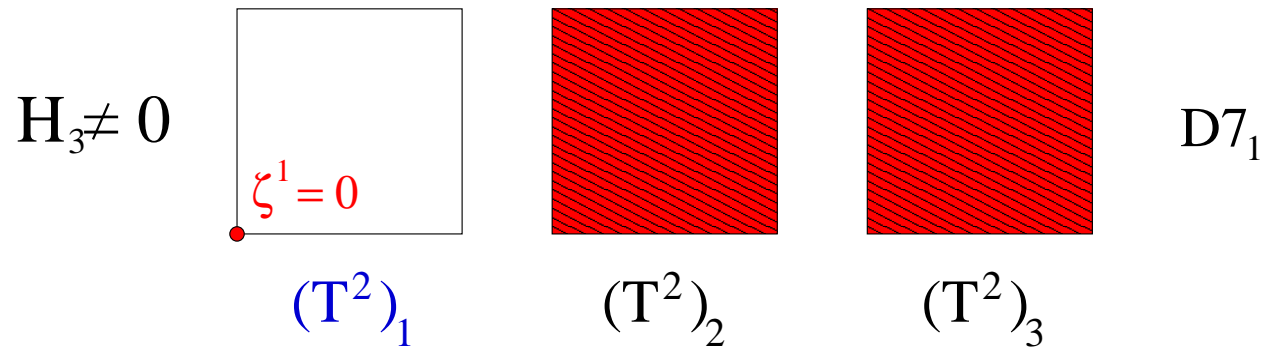
$$\mathcal{F}_{D7_1} = 2\pi\alpha' F_{D7_1} + B_2|_{D7_1}$$

$$\begin{aligned} B|_{D7_1} &= 4\pi^2\alpha' N \left(\zeta_x^1 dx^2 \wedge dx^3 + \zeta_y^1 dy^2 \wedge dx^3 \right) \\ F_{D7_1} &= 2\pi \left(n_1 dx^2 \wedge dx^3 + n_2 dy^2 \wedge dx^3 + \dots \right) \end{aligned}$$

- Additional solutions

$$\begin{aligned} \zeta_x^1 N &= -n_1 \in \mathbb{Z} \\ \zeta_y^1 N &= -n_2 \in \mathbb{Z} \end{aligned}$$

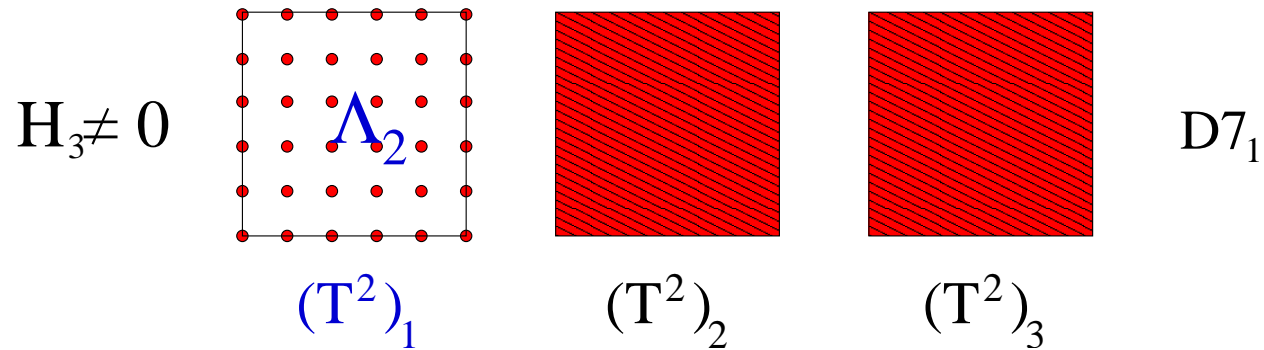
Finding the discretum



$$\mathcal{M}^{G_3}(D7_1) \subset \mathcal{M}(D7_1)$$

$$\{\zeta^1 = 0\} \subset (T^2)_1$$

Finding the discretum



$$\mathcal{M}^{G_3}(D7_1) \subset \mathcal{M}(D7_1)$$

$$\{\zeta^1 = 0\} \subset (T^2)_1$$

↓

$$\Lambda_2 \subset (T^2)_1$$

$$\Lambda_2 = \left\{ \frac{m_x + \tau_1 m_y}{N} \mid m_x, m_y \in \mathbb{Z} \right\} / \Lambda$$

Opening the Landscape

- D7-brane moduli lifting implies a open string reduced moduli space which locally looks like

$$\mathcal{M}^{g_3}(\mathcal{S}_4), \quad \dim_{\mathbb{C}} = h^{0,2} - q$$

Opening the Landscape

- D7-brane moduli lifting implies a open string reduced moduli space which locally looks like

$$\mathcal{M}^{g3}(\mathcal{S}_4), \quad \dim_{\mathbb{C}} = h^{0,2} - q$$

- Globally, however, we expect something of the form

$$\mathcal{M}^{G3}(\mathcal{S}_4) = \Lambda_{2q} \times \mathcal{M}^{g3}(\mathcal{S}_4),$$

Opening the Landscape

- D7-brane moduli lifting implies a open string reduced moduli space which locally looks like

$$\mathcal{M}^{g3}(\mathcal{S}_4), \quad \dim_{\mathbb{C}} = h^{0,2} - q$$

- Globally, however, we expect something of the form

$$\mathcal{M}^{G3}(\mathcal{S}_4) = \Lambda_{2q} \times \mathcal{M}^{g3}(\mathcal{S}_4),$$

- We can actually characterize this new moduli space by

$$\mathcal{M}^{G3}(\mathcal{S}_4) = \left\{ \mathcal{S}_4 \in [\mathcal{S}_4] \mid B_2|_{\mathcal{S}_4} \in A_p^{(1,1)}(\mathcal{S}_4,) \cup H_p^2(\mathcal{S}_4, \mathbb{Z}) \right\},$$

Open String Quantum Numbers

- The spectrum of BPS D7-branes with $U(1)$ bundles is now described in terms of the quantum numbers

i) $[S_4] \in H_4(\mathcal{M}_6, \mathbb{Z}), \quad (\mathcal{S}_4 \text{ hol.}, [H_3|_{\mathcal{S}_4}] = 0)$

ii) $[F] \in H^{(1,1)}(\mathcal{S}_4, \mathbb{R}) \cap H^2(\mathcal{S}_4, \mathbb{Z})$

iii) $\Lambda_{2h^{0,2}}(\mathcal{S}_4)$

Open String Quantum Numbers

• The spectrum of BPS D7-branes with $U(1)$ bundles is now described in terms of the quantum numbers

i) $[S_4] \in H_4(\mathcal{M}_6, \mathbb{Z}), \quad (\mathcal{S}_4 \text{ hol.}, [H_3|_{\mathcal{S}_4}] = 0)$

ii) $[F] \in H^{(1,1)}(\mathcal{S}_4, \mathbb{R}) \cap H^2(\mathcal{S}_4, \mathbb{Z})$

iii) $\Lambda_{2h^{0,2}}(\mathcal{S}_4)$

$i) + ii)$ Considered by [Blumenhagen et al.] in Intersecting D-brane statistics

$iii)$ New

→ Not connected by continuous deformations

→ Can be understood in terms of $B_2|_{\mathcal{S}_4}$ being well-quantized

Conclusions

- The **supersymmetry conditions** for D7-branes in flux compactifications has allowed us understand **moduli lifting** from a **general and geometrical** point of view
- The fact that all (geometrical) moduli are **lifted does not mean** that the D7-brane can be located at **only one point**
 - ⇒ There is a **discretum of D7-brane positions**

Conclusions

- The **supersymmetry conditions** for D7-branes in flux compactifications has allowed us understand **moduli lifting** from a **general and geometrical** point of view
- The fact that all (geometrical) moduli are **lifted does not mean** that the D7-brane can be located at **only one point**
 - ⇒ There is a **discretum of D7-brane positions**
- This defines an **Open String Landscape** directly connected to the **physics of $D = 4$ non-Abelian gauge theories and chiral matter**. Any statistical prediction derived from it should have an interpretation in terms of particle physics

More Conclusions

- This Open String Landscape should affect the counting of type IIB vacua once that we include D7-branes into the story

[Douglas] [Ashok, Douglas] [Denef, Douglas]
[Giryavets, Kachru, Tripathy]

- It should be also taken into account when performing statistics of D-brane configurations in a given background

[Blumenhagen et al.]

- Finally, one should be able to give a general description of this new landscape in terms of F/M-theory

[Lüst, Mayr, Reffert, Stieberger]

[Uranga's Talk]