

# Entanglement Entropy for probe branes.

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work with Han-Chih Chang

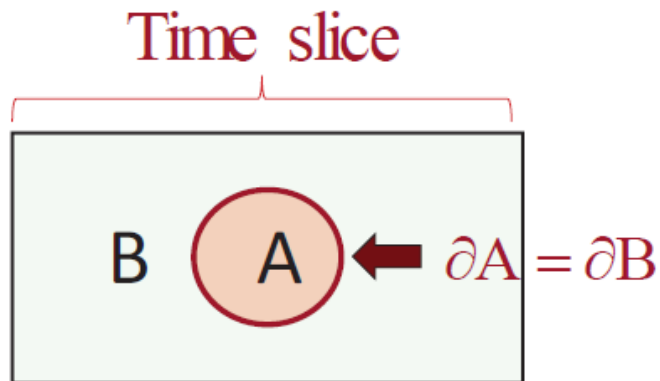
talk at Gauge/Gravity Duality Conference, Munich, Jul 31 2013

# Entanglement Entropy

(1) Spin Chain



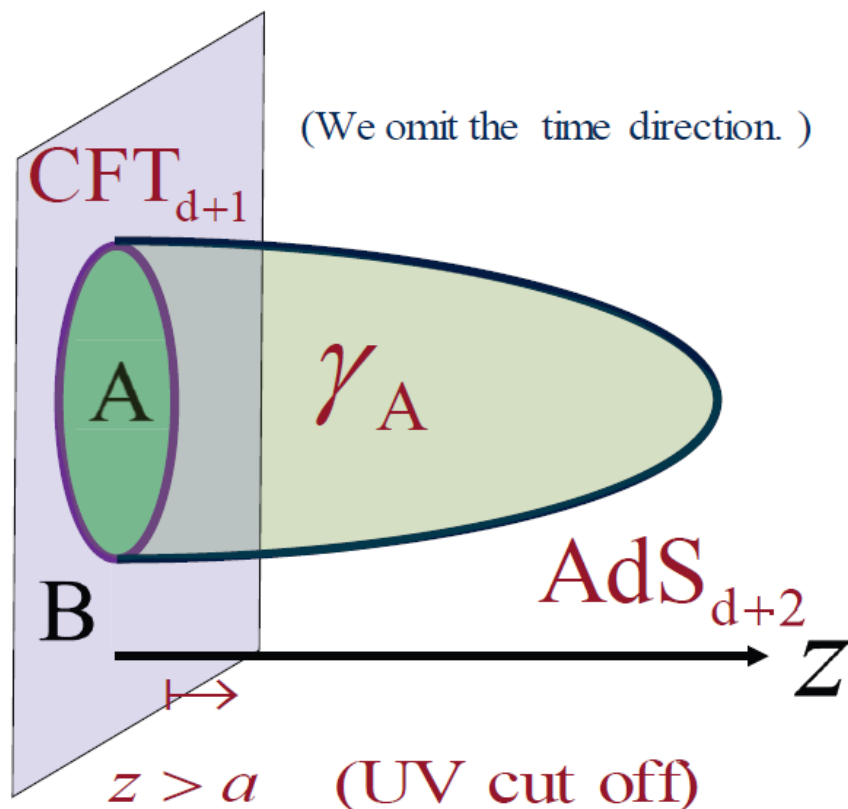
(2) QFT



$$\rho_A = \text{Tr}_B[\rho]$$

$$S_A = -\text{Tr}[\rho_A \log \rho_A]$$

# Holographic Entanglement Entropy

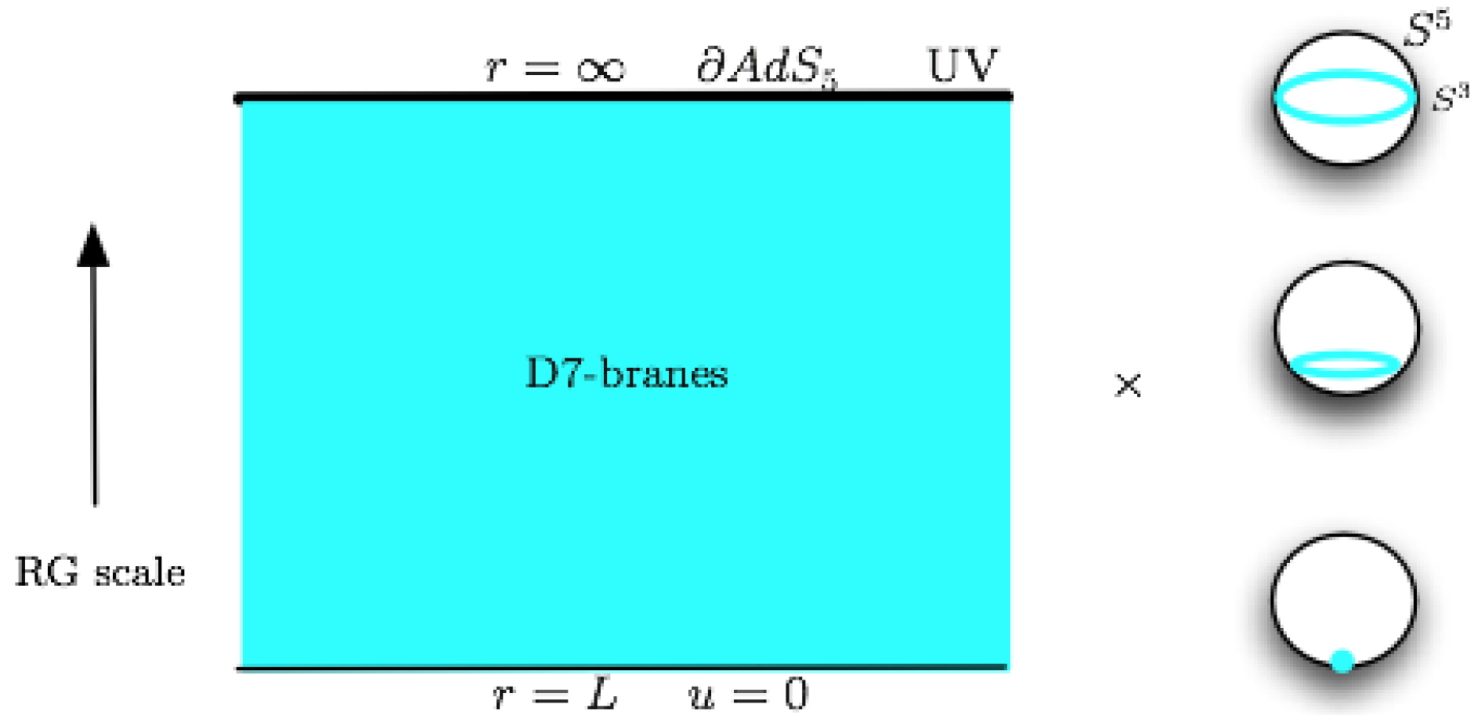


$$S_A = \text{Min}_{\Sigma_A} \left[ \frac{\text{Area}(\Sigma_A)}{4G_N} \right]$$

(Ryu, Takayanagi)

# Probe Branes

(AK, Katz)



(picture from CLMRW-review, 2011)

# Probe Branes

(AK, Katz)

d+1 dimensional  
spacetime



n+1 dimensional  
worldvolume

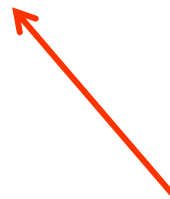


$$\frac{L^{d-1}}{G_N} \gg T_0 L^{n+1} \gg 1$$

Newton's  
constant



Probe brane tension  $N$

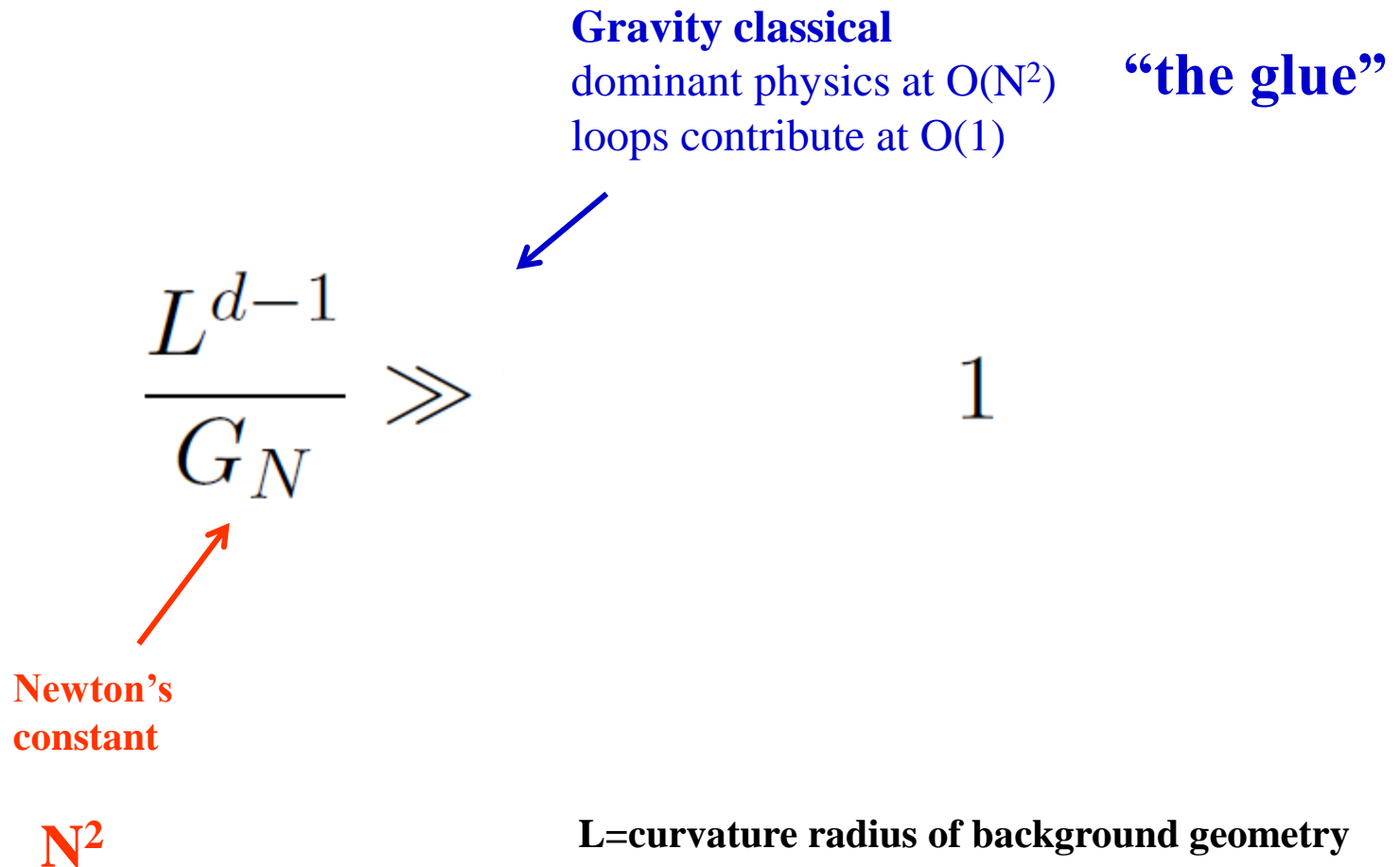


$N^2$

$L$ =curvature radius of background geometry

# Probe Branes

(AK, Katz)



# Probe Branes

(AK, Katz)

**Worldvolume equations classical**  
dominant flavor physics at  $O(N)$   
loops contribute at  $O(1)$

**“the quarks”**

$$T_0 L^{n+1} \gg 1$$

**Probe brane tension  $N$**

**$L$ =curvature radius of background geometry**

# Probe Branes

(AK, Katz)

No gravitational backreaction from probe brane

$O(N^2)$  physics unaffected by order  $N$  physics

$O(N)$  physics: wordvolume of flavor brane and leading order backreaction  
at  $O(1)$  all hell breaks loose: 2<sup>nd</sup> order backreaction, glue loops, quark loops

$$\frac{L^{d-1}}{G_N} \gg T_0 L^{n+1} \gg 1$$

Newton's  
constant

$N^2$

Probe brane tension  $N$

$L$ =curvature radius of background geometry



# The quark sector

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Gravity bad -- DBI good.

(even I can solve it)

How far can I get without ever having to solve Einstein's equations?

There are some new quantities in the quark sector, the flavor sector that are manifestly only due to DBI fields (e.g. chiral condensate) <sup>9</sup>

# Gravity bad – DBI good

(see e.g. AK, O'Bannon, Thompson)

Generically, at order  $N$  backreaction matters.

$$S = N^2 S_{grav} + N S_{probe} \quad \text{glue and quark sector}$$

backreaction suppressed....  $\delta h \propto G_N T_{\mu\nu} \propto 1/N h$

$$\varepsilon = N^2 \varepsilon_{grav} + N \left( \varepsilon_{probe} + \frac{\delta \varepsilon_{grav}}{\delta g} h \right)$$

$\varepsilon = \delta S / \delta g_{00}$  = energy density

... but enters at order  $N$

# Gravity bad – DBI good

(see e.g. AK, O'Bannon, Thompson)

Exception: the free energy = on-shell action.

$$S = N^2 S_{grav} + N S_{probe} \quad \text{glue and quark sector}$$

backreaction suppressed

$$\delta h \propto G_N T_{\mu\nu} \propto 1/N \quad h$$

$$\omega = N^2 S_{grav} + N \left( S_{probe} + \frac{\delta S_{grav}}{\delta g} h \right) \quad =0$$

contribution due to backreaction vanishes due to equations of motion!

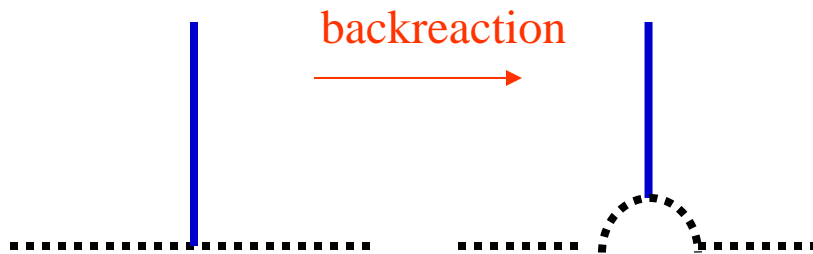
# Gravity bad – DBI good

(see e.g. AK, O’Bannon, Thompson)

$$\omega = N^2 S_{grav} + N S_{probe}$$

**good news! all equilibrium properties can be calculated without ever having to deal with backreaction!!!!**

Including thermal entropy:



or

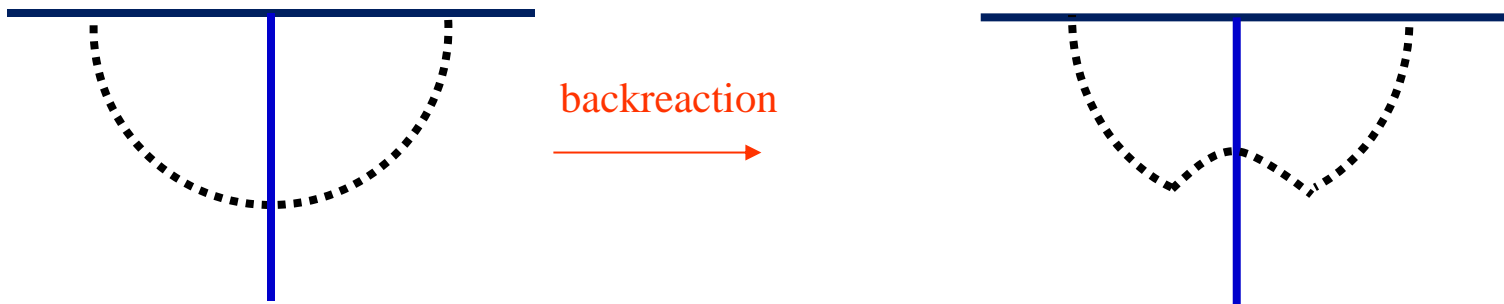
$$S = - \frac{\partial \omega}{\partial T}$$

but “cheat” works!

**honest calculation:  
calculate change in horizon area**

# DBI bad – Gravity good!!!

But for the EE we are stuck with needing leading order backreaction to even get order N contribution.



$$S^{EE} = \underbrace{S_{grav}^{EE}}_{\mathcal{O}(N^2)} + \frac{\delta Area}{\underbrace{4G}_{\mathcal{O}(N)}} \quad \begin{array}{l} \leftarrow \mathcal{O}(1/N) \\ \leftarrow \mathcal{O}(N^2) \end{array}$$

In this work we will give a general expression for  $\delta Area$

# Exceptions:

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Casini, Huerta, Myers:

- spherical entangling surface
- conformal field theory

EE

conformal map



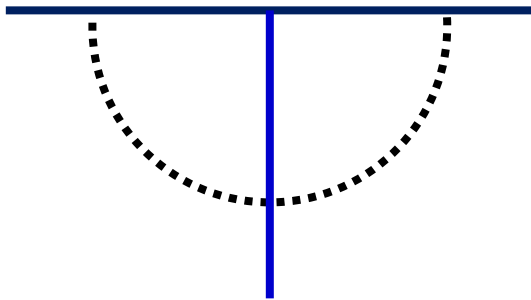
thermal entropy  
on hyperboloid

Jensen, O'Bannon:

still applies for conformal flavors  
(massless quarks!) on conformal defect

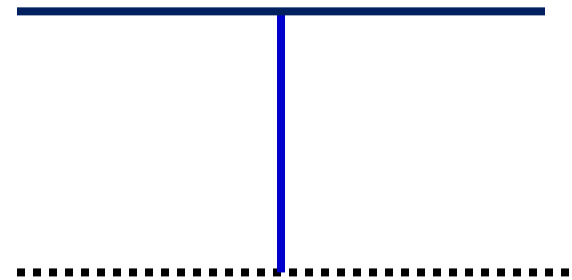
In this case cheat still applies.

# Jensen-O'Bannon Calculation



(CHM)

=



$$ds^2 = \frac{-dt^2 + dx^2 + dz^2}{z^2}$$

**T=0**

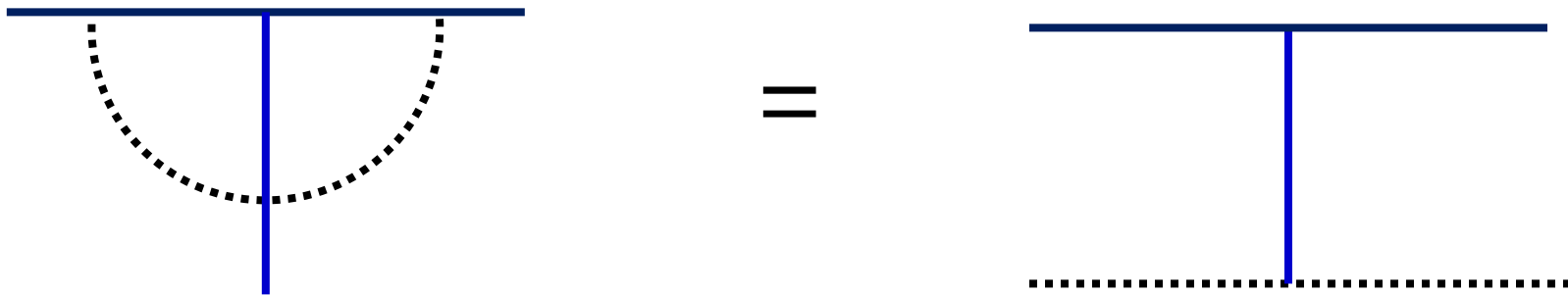
$$ds^2 = -h(r)dt^2 + dH^2 + dr^2/h(r)$$

$$h(r) = r^2 - 1$$

**T=1/2π**

Casimir stress tensor on H appears thermal  
Entangling surface = boundary of H

# Jensen-O'Bannon Calculation



$O(N^2)$ : (CHM)

$$S = \frac{V_{d-1}^H}{4G_N}$$

$\frac{d}{dT}$

←

$r_H = 2\pi T / (d - 1)$

constant

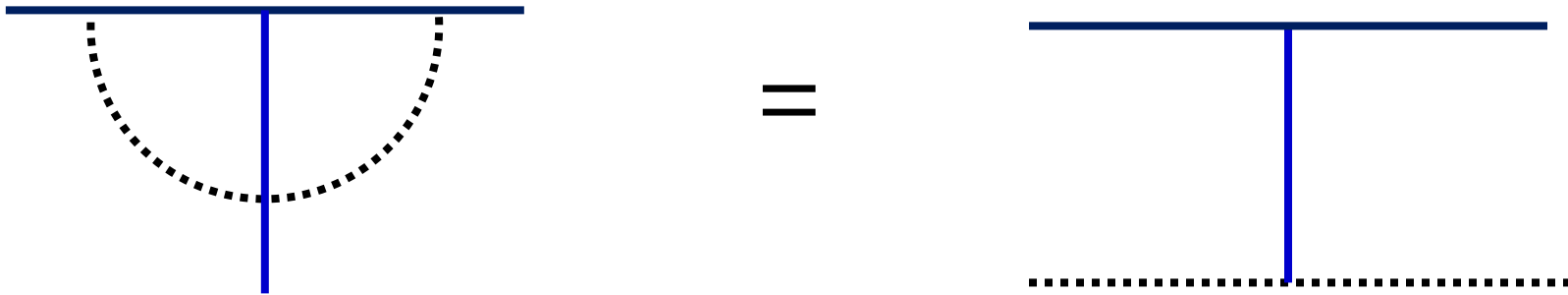
$$S = \frac{1}{16\pi G_N} \int d^{d+1}x \sqrt{g} \left( R + \frac{d(d-1)}{L^2} \right)$$

**Free energy =  $C_0$  volume of spacetime**  
**=  $-(C_0/d) r_h^d \text{ vol}(H)$**

$$C_0 = d / (8\pi G_N)$$



# Jensen-O'Bannon Calculation



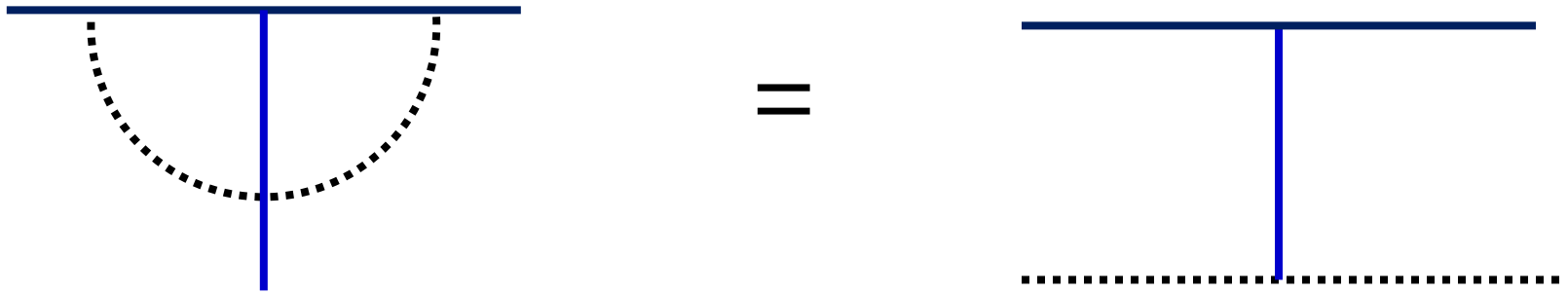
$O(N^2)$ : (CHM)

$$S = \frac{V_{d-1}^H}{4G_N}$$

$$\begin{aligned}
 V_m^H &= V_{m-1}^S \int_{\epsilon}^1 dy \frac{(1-y^2)^{(m-2)/2}}{y^m} \\
 &= p_1 \left(\frac{1}{\epsilon}\right)^{m-1} + p_3 \left(\frac{1}{\epsilon}\right)^{m-3} + \dots \\
 &\dots + \begin{cases} p_{m-1} \left(\frac{1}{\epsilon}\right) + p_m + \mathcal{O}(\epsilon) & m \text{ even} \\ p_{m-2} \left(\frac{1}{\epsilon}\right) + q \log(\epsilon) + \mathcal{O}(1) & m \text{ odd.} \end{cases}
 \end{aligned}$$

area law

# Jensen-O'Bannon Calculation



**O(N):** (JO)

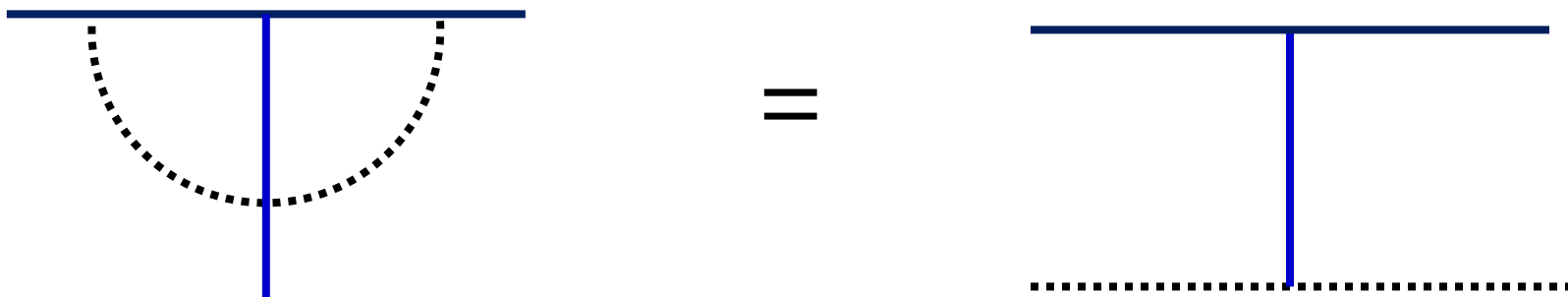
$$S = \frac{2\pi T_0}{d-1} V_{n-1}^H \leftarrow \frac{d}{dT}$$

$$S_{probe} = -T_0 \int d^{n+1}z \sqrt{g_I}$$

**Free energy ~ volume of spacetime**

(for  $n < d$ ,  $1/(d-1)$  replaced by  $1/d$  for  $n=d$ )

# Jensen-O'Bannon Calculation



**O(N):** (JO)

$$S = \frac{2\pi T_0}{d-1} \underline{V_{n-1}^H}$$

“flavor central charge”  
number of DOFs on defect

Has exactly the functional form  
of a spherical entangling surface  
in an n+1 dimensional CFT.

# Limitations of Jensen – O'Bannon

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- Massless flavors only

**no topological phases**

**not applicable to probe branes realizing  
quantum Hall effect, topological insulators, ....**

- No worldvolume gauge fields

**no novel compressible quantum liquids**

**no holographic quantum liquid**

# Limitations of Jensen – O'Bannon

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- Massless flavors only
- No worldvolume gauge fields

**not applicable to any of the systems we really want to be able to calculate the EE in!**

Want a method that works for any probe brane

**use Jensen-O'Bannon to validate our new method.**

# EE for probe branes.

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$$S_A = (\pi T_0) \int (d^{d-1} w \sqrt{\gamma}) (d^{n+1} z \sqrt{g_I}) \left( T_{min}^{\mu\nu} G_{\mu\nu\rho\sigma} T_{probe}^{\rho\sigma} \right)$$

tension  
 $O(N)$  constant

minimal  
area

probe  
worldvolume

double integral

gravitational  
Green's function

# EE for probe branes.

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= gravitational interaction between two energy distributions

$$S_A = (\pi T_0) \int (d^{d-1}w \sqrt{\gamma}) (d^{n+1}z \sqrt{g_I}) \left( T_{min}^{\mu\nu} G_{\mu\nu\rho\sigma} T_{probe}^{\rho\sigma} \right)$$

$$S_{min} = \frac{1}{4G_N} \int d^{d-1}w \sqrt{\gamma} \equiv \frac{1}{4G_N} \int d^{d-1}w \mathcal{L}_{min}$$

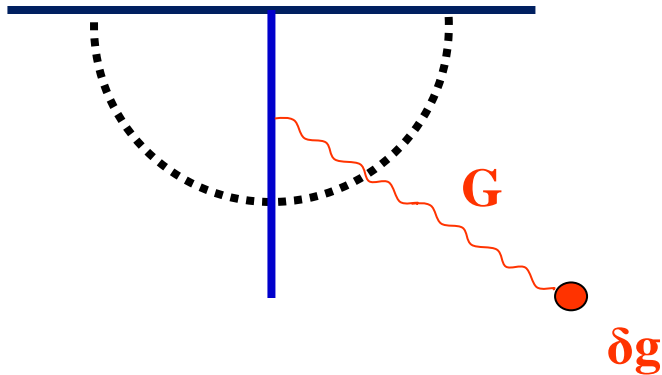
$$T_{min}^{\mu\nu} = \frac{2}{\sqrt{-g}} \frac{\delta(\sqrt{-g} \mathcal{L}_{min})}{\delta g_{\mu\nu}}$$

“stress tensor” of  
minimal area

“stress tensor” of  
probe brane

# EE for probe branes - derivation

$$S_A = (\pi T_0) \int (d^{d-1}w \sqrt{\gamma}) (d^{n+1}z \sqrt{g_I}) \left( T_{min}^{\mu\nu} G_{\mu\nu\rho\sigma} T_{probe}^{\rho\sigma} \right)$$



$\delta g_{\mu\nu}$   
linearized backreaction



# EE for probe branes - derivation

$$S_A = (\pi T_0) \int (d^{d-1}w \sqrt{\gamma}) (d^{n+1}z \sqrt{g_I}) \left( T_{min}^{\mu\nu} G_{\mu\nu\rho\sigma} T_{probe}^{\rho\sigma} \right)$$

corresponding change in minimal area:

$$\delta S_{min} = \frac{1}{4G_N} \int d^{d-1}w \sqrt{\gamma} \left( \frac{T_{min}^{\mu\nu}}{2} (\delta g)_{\mu\nu} + \frac{\delta \mathcal{L}_{min}}{\delta x_M^\mu} \delta x_M^\mu \right)$$

=0 by eom.

change of  
embedding

# Comments

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- Works in time-dependent backgrounds

Hubeny, Rangamani, Takayanagi:

Minimal area  $\rightarrow$  Extremal area

same action!

use retarded Green's function for G

# Comments

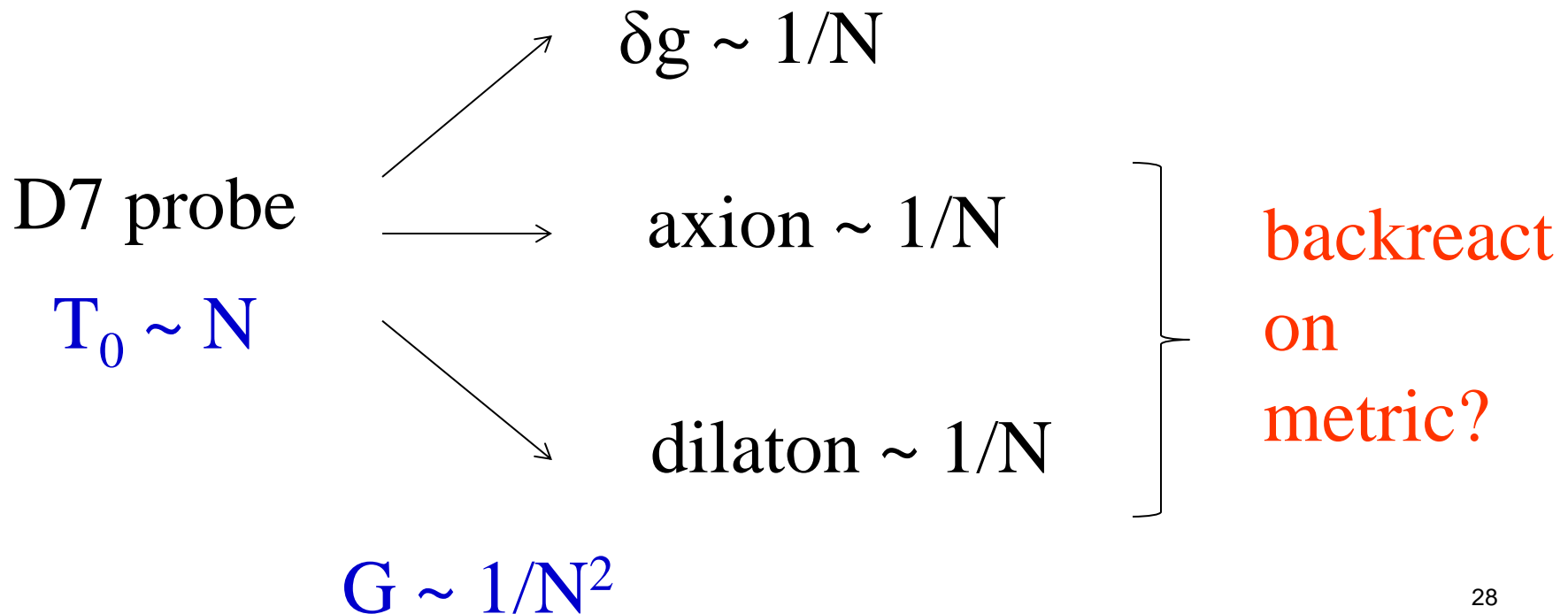
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- Works with higher curvature corrections
  - minimal area requirement gets replaced with more general surface
  - still follows from an action principle, action defining the surface is area + curvature corrections
  - still allows definition of  $T^{\min}_{\mu\nu}$

# Properties

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No secondary backreaction!



# Secondary Backreaction

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$$\text{axion} \sim 1/N \quad G \sim 1/N^2$$

$$\text{IIB-action} = 1/G (\text{axion})^2$$

$$\text{stress tensor} = N^2 (\text{axion})^2 \sim O(1)$$

$$\delta g \sim G (\text{stress tensor}) \sim 1/N^2$$

secondary backreaction is subleading

Exception:

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turned on in background



$$\text{axion} \sim 1/N$$

$$\text{Axion} \sim 1$$

$$G \sim 1/N^2$$

$$\text{IIB-action} = 1/G (\text{Axion} + \text{axion})^2$$

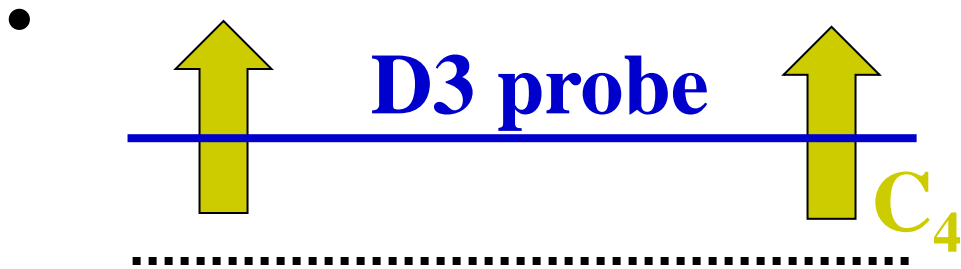
$$\text{stress tensor} = \dots + N^2 (\text{axion Axion}) \sim O(N)$$

$$\delta g \sim G (\text{stress tensor}) \sim 1/N$$

secondary backreaction is important if brane sources field that has non-trivial background

# Exception: Examples

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- D6 flavor brane in ABJM  
(ABJM has RR 2-form background)

**But: D5 with  $F_{rt}$  is fine. Sources  $C_4$ ,  
but orthogonal to background.**

# Properties

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Backreaction in internal space irrelevant.

trace reversed Einstein:

$$R_{\mu\nu} = \tilde{T}_{\mu\nu} \quad \tilde{T}_{\mu\nu} = T_{\mu\nu} - \frac{1}{D-2} g_{\mu\nu} T_{\rho}^{\rho}$$

Entangling surface is:

- codimension 2 minimal surface
- wrapping internal manifold

**makes co-dimension 2 special**

and hence does not source internal tensor modes




# Properties

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Backreaction in internal space irrelevant.

Instead of integrating over D-dim product space, can do integral in (d+1) dimensional spacetime with “effective” probe stress tensor.

$$T_{\mu\nu}^{probe,eff} = \int_{x_I} \sqrt{g_I} T_{\mu\nu}^{probe}$$

  
**internal space**

# No internal $O(N)$ backreaction:

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EE for D3/D7 ( $\text{AdS}_5 \times S^3$ ) =

EE for spacetime filling probe in  $\text{AdS}_5$

EE for D3/D5 ( $\text{AdS}_4 \times S^2$ ) =

EE for co-dim 1 probe in  $\text{AdS}_5$  =

Randall-Sundrum brane

← very simple to get  
fully backreacted  
metric!

# Discussion

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Gave universal formula for probe EE. Two important properties:

- no secondary backreaction
- no internal backreaction

# Discussion

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We checked our formula in the simplest examples:

- spherical entangling surface
- strip entangling surface

for

- D3/D7
- D3/D5
- toy models (spacetime filling and RS)

and found perfect agreement with

- fully backreacted solution (toy-models)
- Jensen-O'Bannon calculation (sphere)

# To-Do list:

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A whole new world to explore, including:

**topological phases**

**probe branes realizing  
quantum Hall effect, topological insulators, ....**

**novel compressible quantum liquids**

**holographic quantum liquid,  
 $\theta=d-1$  non-relativistic critical points**